

Stability in and of de Sitter space

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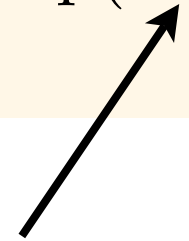
Outline

- Review of instantons and exponential decay
- The asymmetric double well
- Field theory in flat space & finite volume
- Coleman - De Luccia
- Canonical quantization methods and the failure of the probe approximation
- Outlook

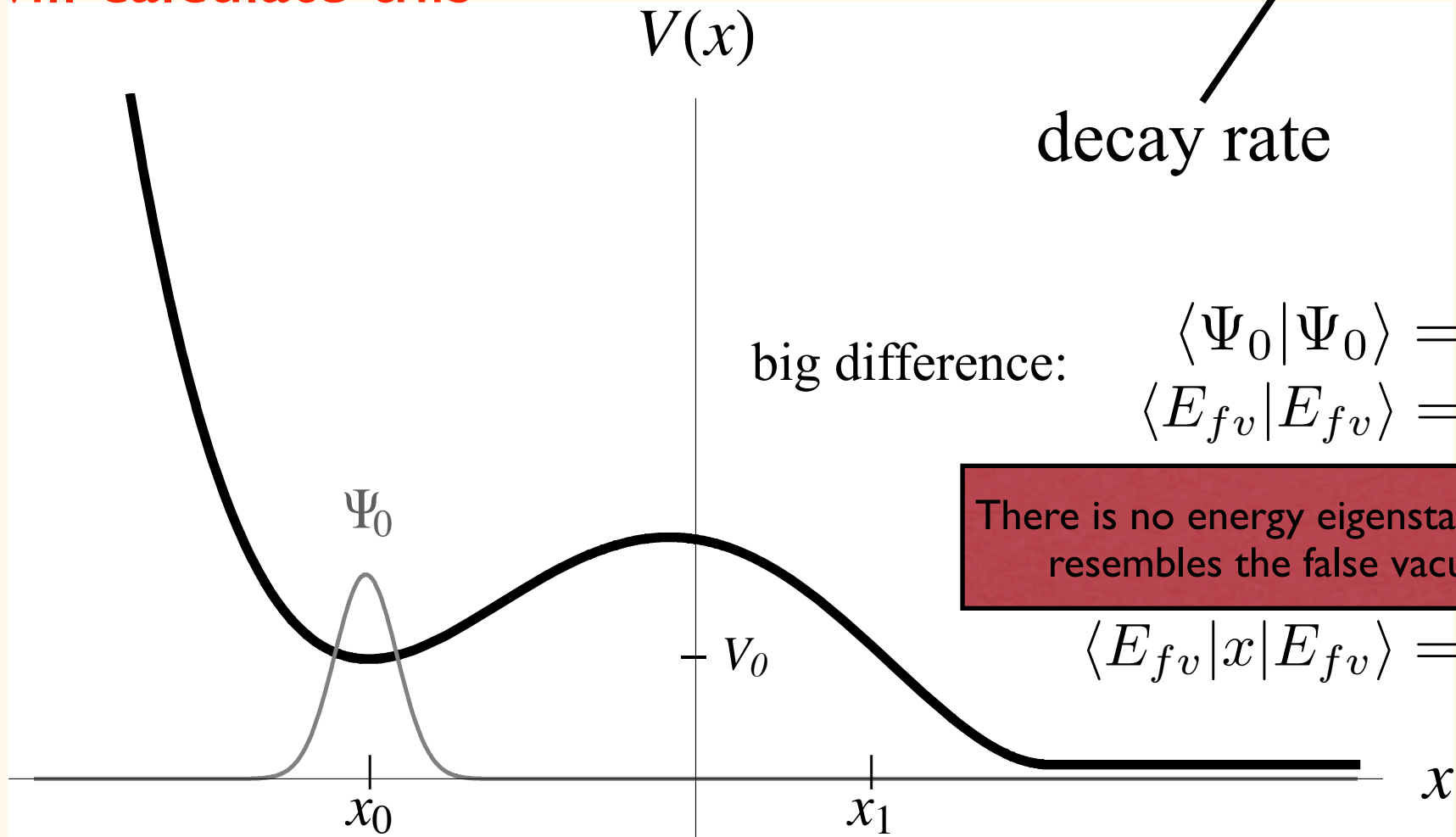
exponential decay

$$|\langle \Psi_0 | \exp(-iHt) | \Psi_0 \rangle|^2 \approx \exp(-\Gamma|t|)$$

We will calculate this

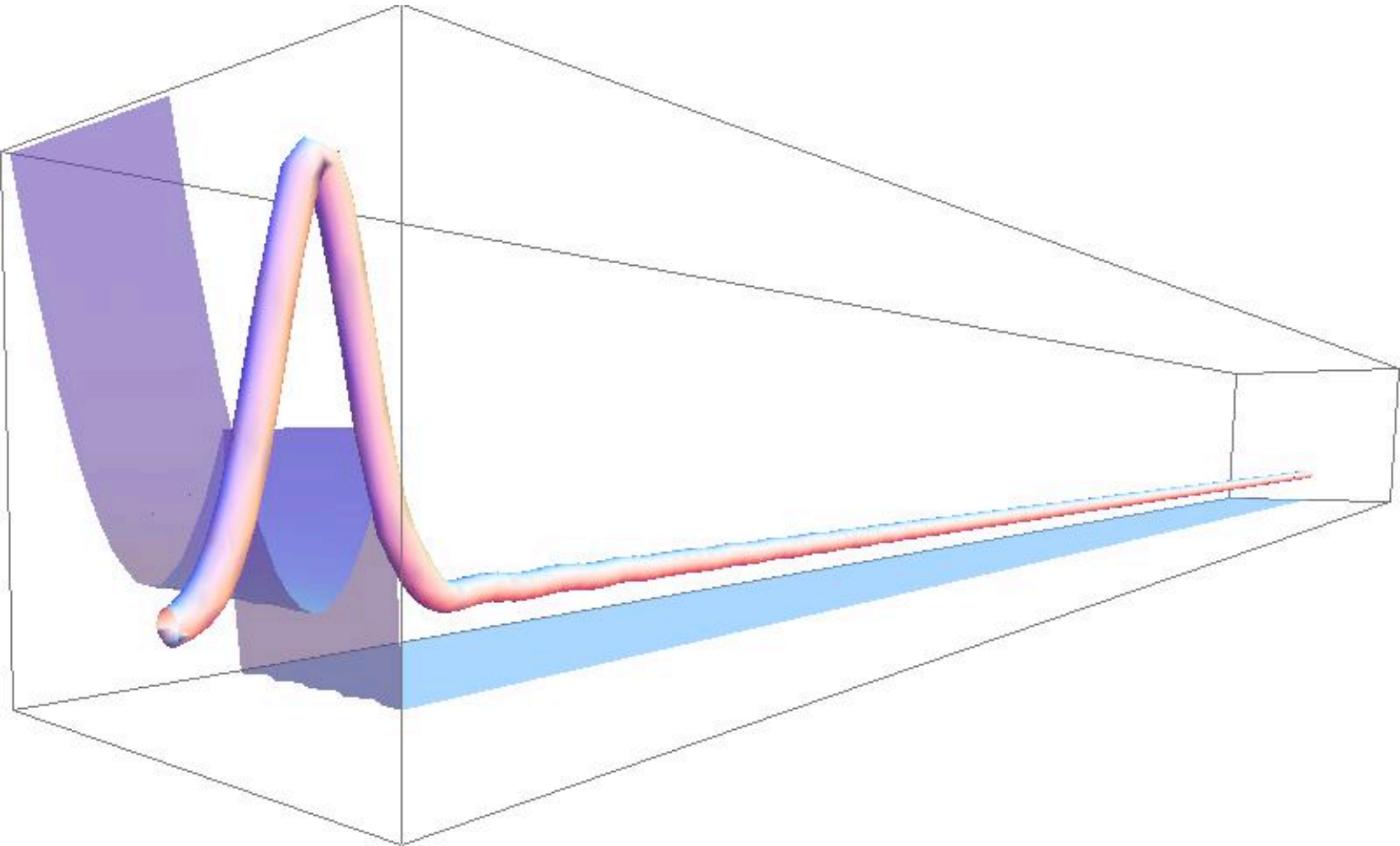


decay rate

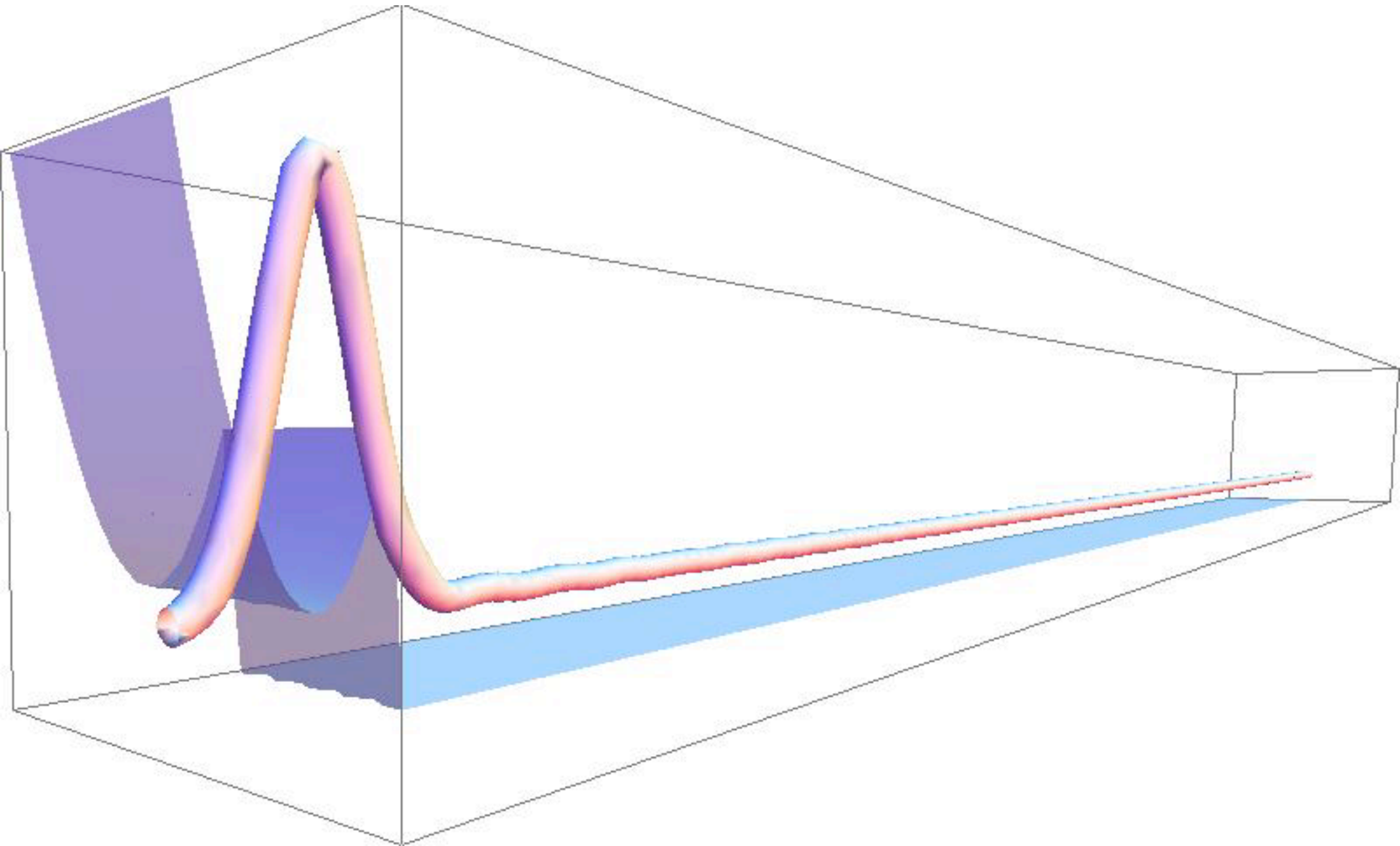


“false vacuum” $|\Psi_0\rangle =$ ground state of *perturbative* Hamiltonian

Early time behavior: Probability current flows outward.



Early time behavior: Probability current flows outward.



Late time behavior: Survival probability is zero. (not pictured)

instanton method “The uses of instantons”

$$|\langle \Psi_0 | \exp(-iHt) | \Psi_0 \rangle|^2 \approx \exp(-\Gamma|t|)$$

$$\langle x_f | e^{-iHt} | x_i \rangle = \int [dx] e^{iS[x]}$$

$t \rightarrow -it = -T$ This is the *definition* of the path integral
(consistent with Feynman’s pole prescription)

$$S_E = \int_{-T/2}^{T/2} \left(\frac{1}{2} \dot{x}^2 + V(x) \right) d\tau$$

$$V(x) \rightarrow -V(x)$$

$$\langle x_0 | e^{-HT} | x_0 \rangle = \int [dx] e^{-S_E[x]}$$

On the left: Large T picks out the lowest energy states

On the right: We will use method of steepest descent to calculate the late time behavior of the low energy states.

$$H = H_{pert.} + \Delta$$

Assume no degeneracies in $H_{pert.}$ (no resonances)

$$\langle x_0 | e^{-HT} | x_0 \rangle = \sum_n \langle x_0 | e^{-(H_{pert.} + \Delta)T} | n \rangle \langle n | x_0 \rangle$$

H is
unbounded below

$$\xrightarrow{\text{large } T} e^{-(E_{fv} + \delta_{fv})T} |\langle x_0 | \Psi_0 \rangle|^2$$

If we treat Ψ_0 like an
eigenstate of H , its energy
has an imaginary part.

Actually, Δ has off-
diagonal terms which
preserve unitarity.

$$\delta_{fv} = -\frac{i}{2}\Gamma$$

Degeneracies mean $H_{pert.}$ basis has
arbitrariness which may be totally
destroyed by effects of Δ .

non-perturbative correction
to false-vacuum "energy"

$$|\langle \Psi_0 | \exp(-iHt) | \Psi_0 \rangle|^2 \approx \exp(-\Gamma|t|)$$

Method of steepest descent

$$S_E[x] \approx \sum_{x_{cl}} S_E[x_{cl}] + \frac{1}{2} \delta x \left. \frac{\delta^2 S_E}{\delta x^2} \right|_{x_{cl}} \delta x$$

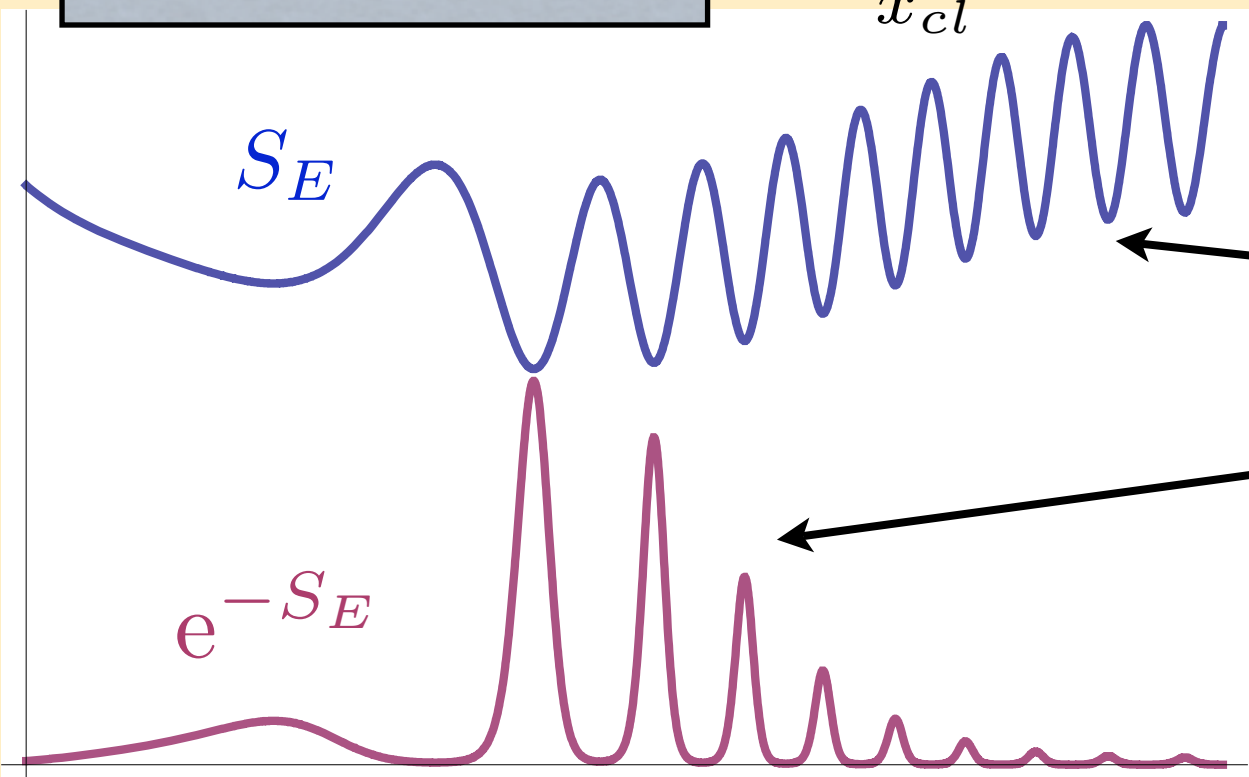
$$\left. \frac{\delta^2 S_E}{\delta x^2} \right|_{x_{cl}} = -\partial_\tau^2 + V''(x_{cl}(\tau))$$

Fluctuations
= **measure**

expand about the local minima of $S_E[x]$

$$\int [dx] e^{-S_E[x]} \approx \sum_{x_{cl}} e^{-S_E[x_{cl}]}$$

$$\sqrt{\frac{1}{\det \left[\frac{\delta^2 S_E}{\delta x^2} \right]}}$$



many minima

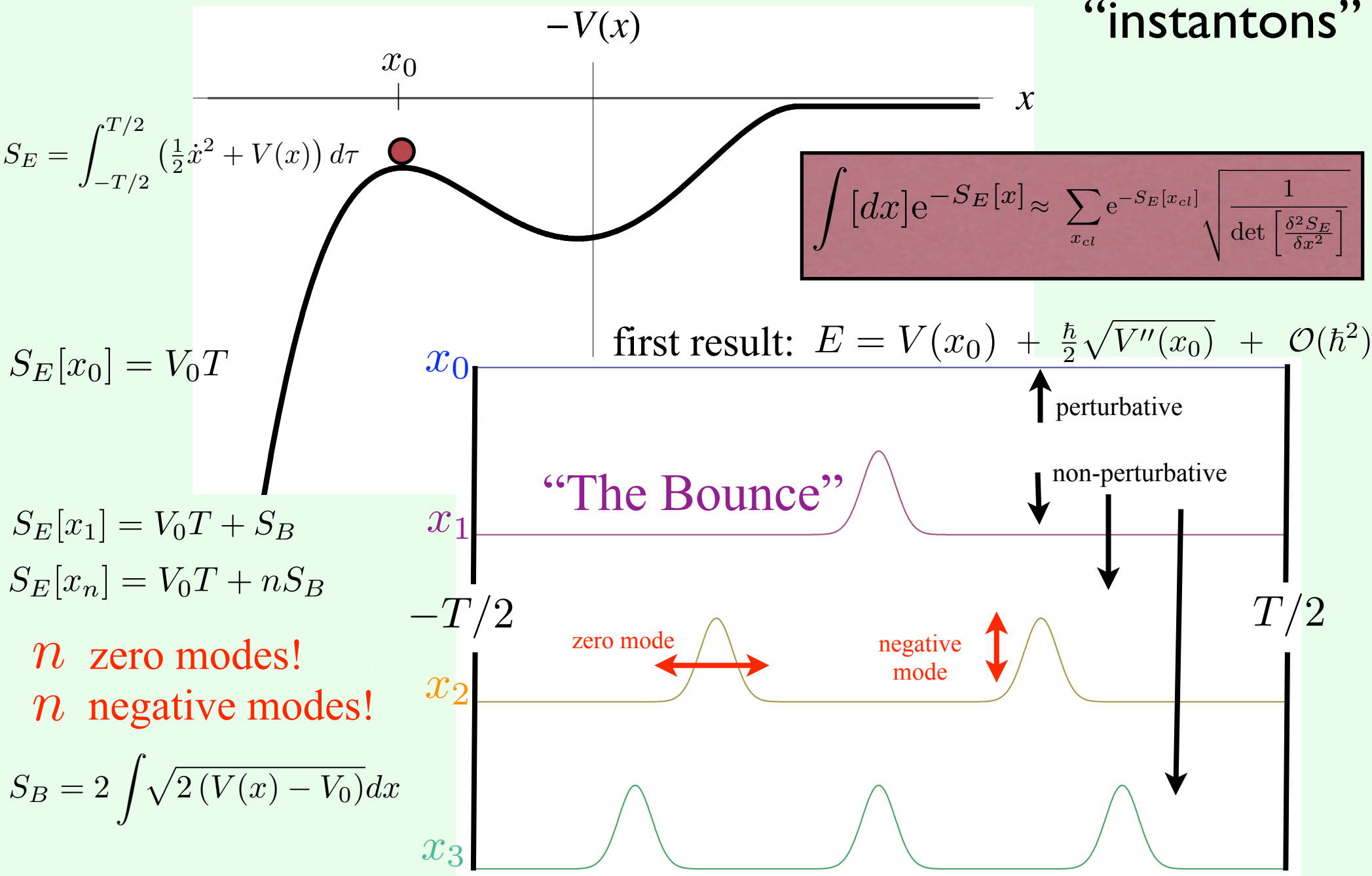
many Gaussians

$$\int_{-\infty}^{\infty} \frac{dx}{\sqrt{\pi}} e^{-\lambda x^2} = \frac{1}{\sqrt{\lambda}}$$

What are the x_{cl} ?

any number of bounces ≥ 0 .

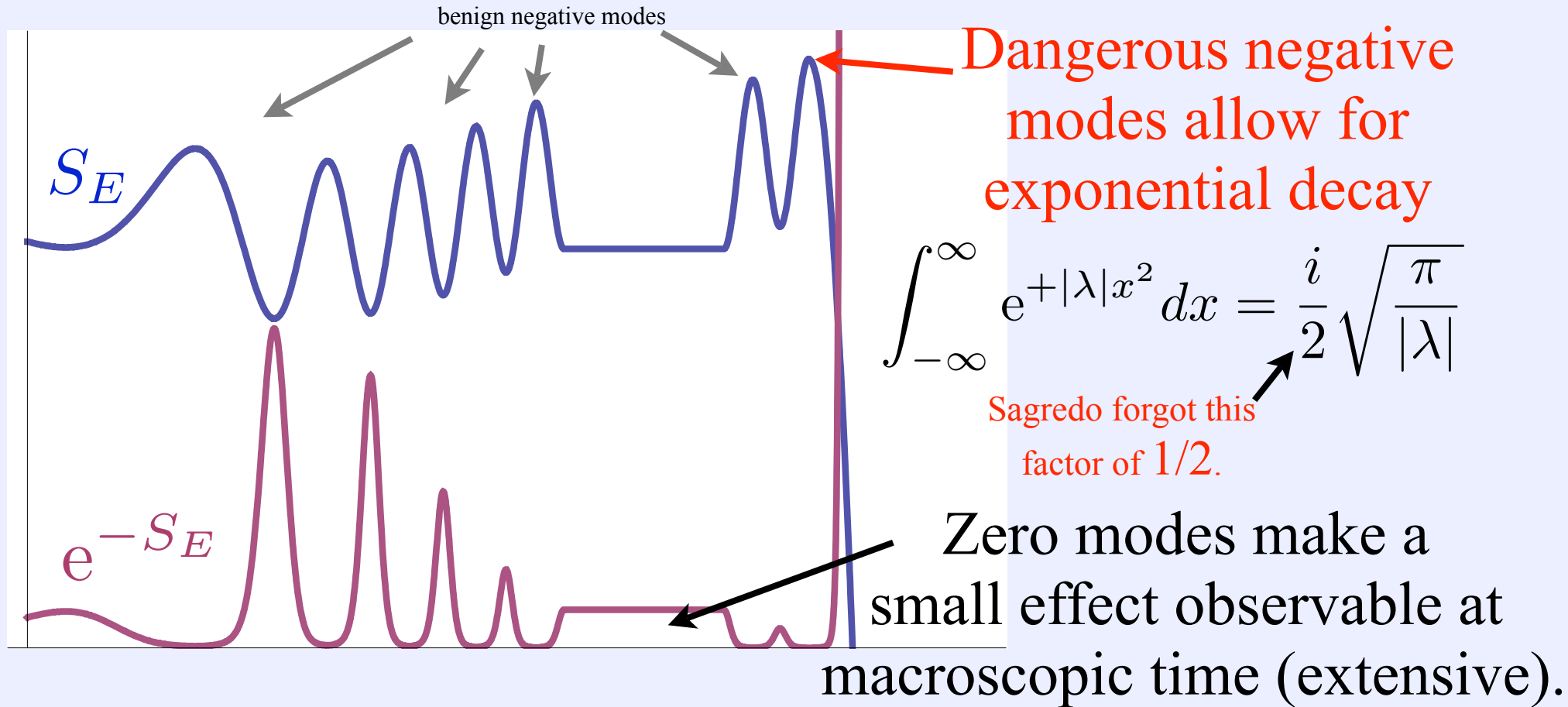
“instantons”



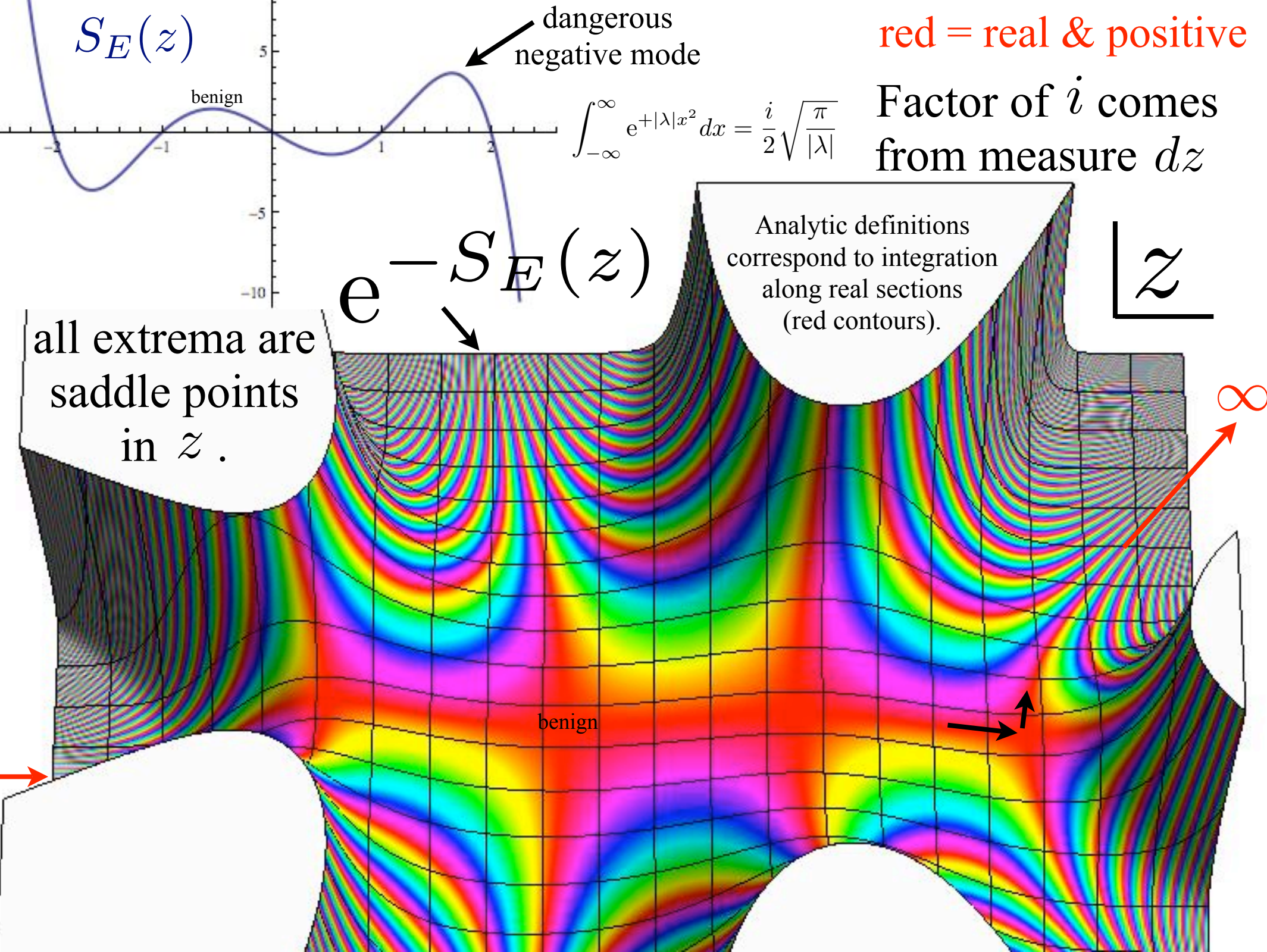
Each zero mode contributes a large, but finite factor T .

$$\int_{-T/2}^{T/2} e^{-0\tau^2} d\tau = T$$

Each “dangerous” negative mode contributes a *tiny* imaginary factor.



When a physical quantity appears divergent, it is defined via analytic continuation. We must choose amongst the various branches by appealing to the physics. In the case at hand, the “unstable state” is defined as the ground state of an analytically deformed potential. In the figure above, every maxima marks a crossroad. Only the one leading to an actual divergence must be avoided.



$x_n \equiv n$ widely separated bounces

$$S_E[x_n] = S_E[x_0] + nS_B \quad S_B = 2 \int \sqrt{2(V(x) - V_0)} dx$$

$$\det \left[\frac{\delta^2 S_E}{\delta x^2} \right]_{x_n}^{-\frac{1}{2}} = \frac{T^n}{n!} \left(\frac{-i}{2} \right)^n \det' \left[\left| \frac{\delta^2 S_E}{\delta x^2} \right| \right]_{x_n}^{-\frac{1}{2}} = \frac{T^n}{n!} \left(\frac{-i}{2} \right)^n \det' \left[\left| \frac{\delta^2 S_E}{\delta x^2} \right| \right]_{x_B}^{\frac{n}{2}} \det \left[\frac{\delta^2 S_E}{\delta x^2} \right]_{x_0}^{-\frac{1}{2}}$$

$$\sqrt{\frac{1}{\det \left[\frac{\delta^2 S_E}{\delta x^2} \right]_{x_0}}} = e^{-\frac{1}{2}\omega T + \mathcal{O}(\hbar^2)T} \quad \omega = \sqrt{V''(x_0)}$$

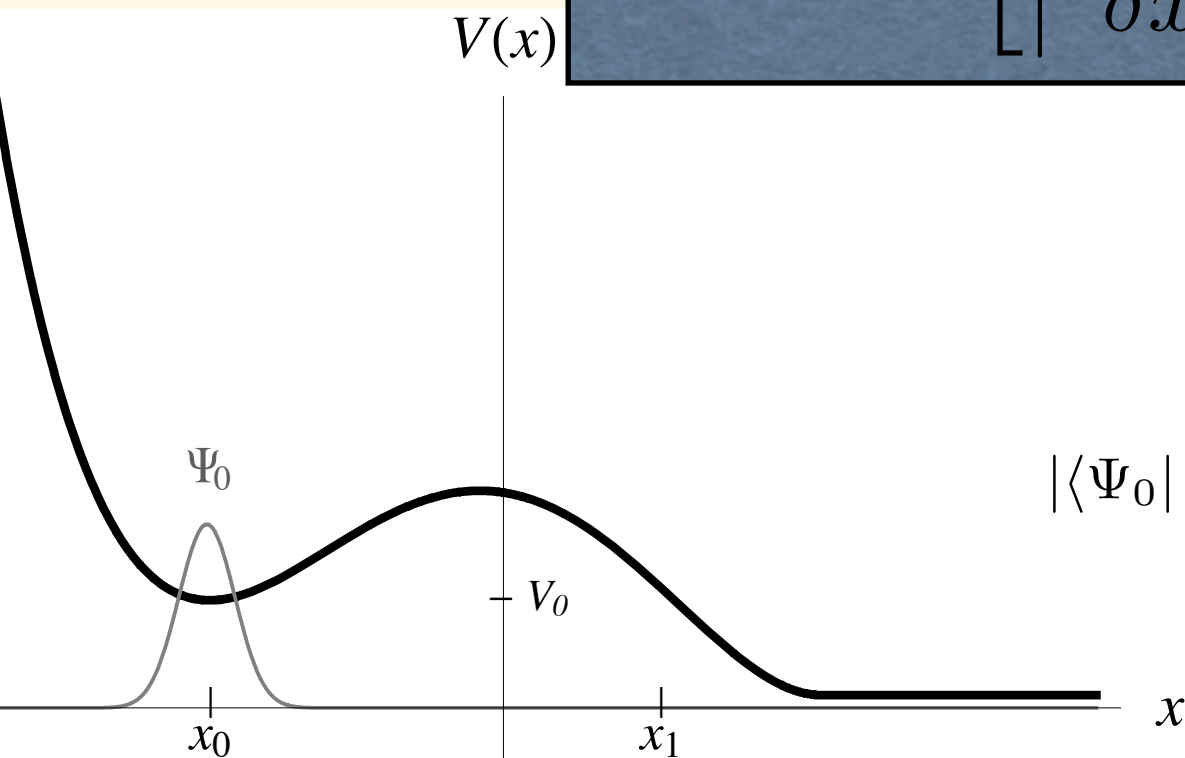
$$\begin{aligned} \int [dx] e^{-S_E[x]} &\approx \sum_{x_{cl}} e^{-S_E[x_{cl}]} \sqrt{\frac{1}{\det \left[\frac{\delta^2 S_E}{\delta x^2} \right]}} \\ &= \sum_{n=0}^{\infty} e^{-(V_0 + \frac{\omega}{2} + \mathcal{O}(\hbar^2))T} \left(\frac{T^n}{n!} \left(\frac{-i}{2} \right)^n \det' \left[\left| \frac{\delta^2 S_E}{\delta x^2} \right| \right]_{x_B}^{\frac{n}{2}} e^{-nS_B} \right) \\ &= e^{-\left(V_0 + \frac{\omega}{2} - \frac{i}{2} \det' \left[\left| \frac{\delta^2 S_E}{\delta x^2} \right| \right]_{x_B}^{-\frac{1}{2}} e^{-S_B} \right) T} \end{aligned}$$

At large T

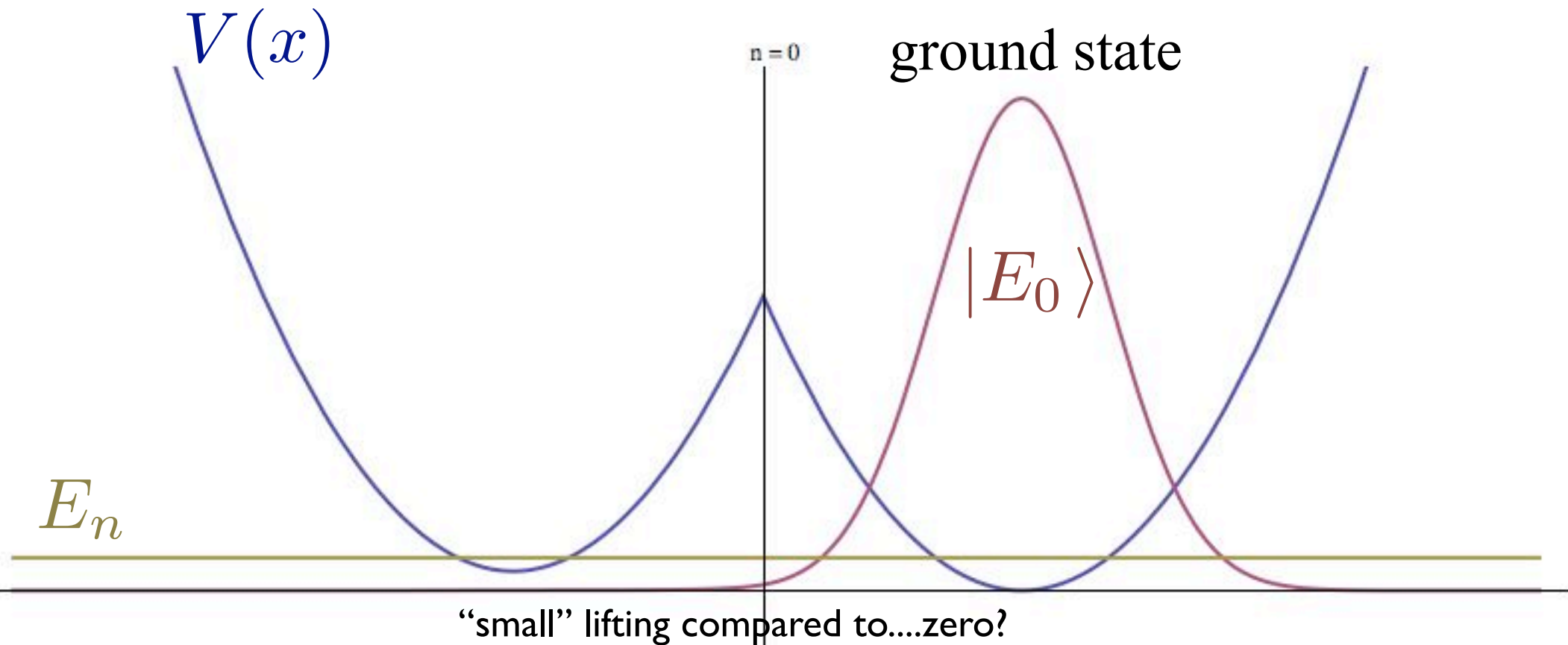
$$e^{-(E_{fv} + \delta_{fv})T} = e^{-\left(V_0 + \frac{\omega}{2} - \frac{i}{2} \det' \left[\left| \frac{\delta^2 S_E}{\delta x^2} \right| \right]_{x_B}^{-\frac{1}{2}} e^{-S_B} \right) T}$$

$$\delta_{fv} = -\frac{i}{2} \Gamma$$

$$\Gamma = \det' \left[\left| \frac{\delta^2 S_E}{\delta x^2} \right| \right]_{x_B}^{-\frac{1}{2}} e^{-S_B}$$



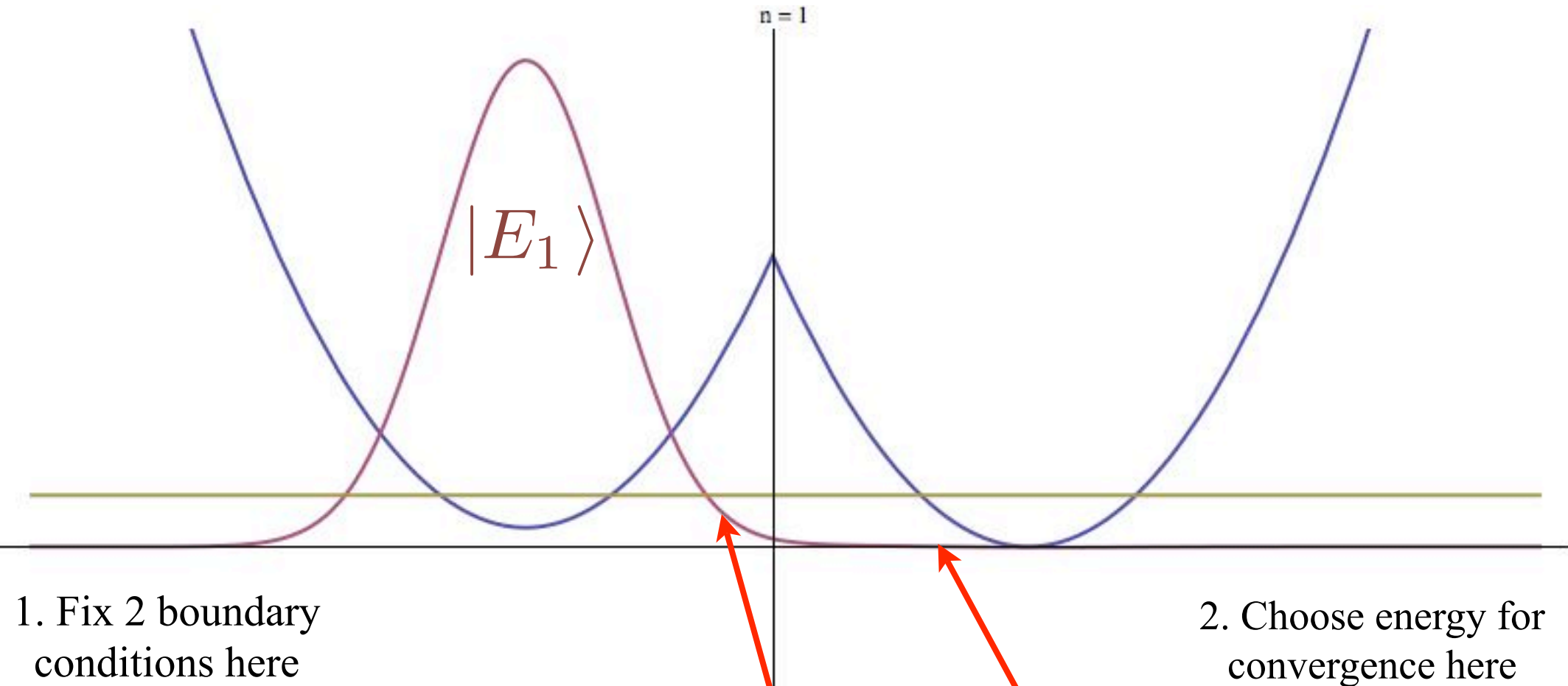
$$|\langle \Psi_0 | \exp(-iHt) | \Psi_0 \rangle|^2 \approx \exp(-\Gamma|t|)$$



The asymmetric double well

not a perturbation of the symmetric double well!

A Stable False Vacuum!

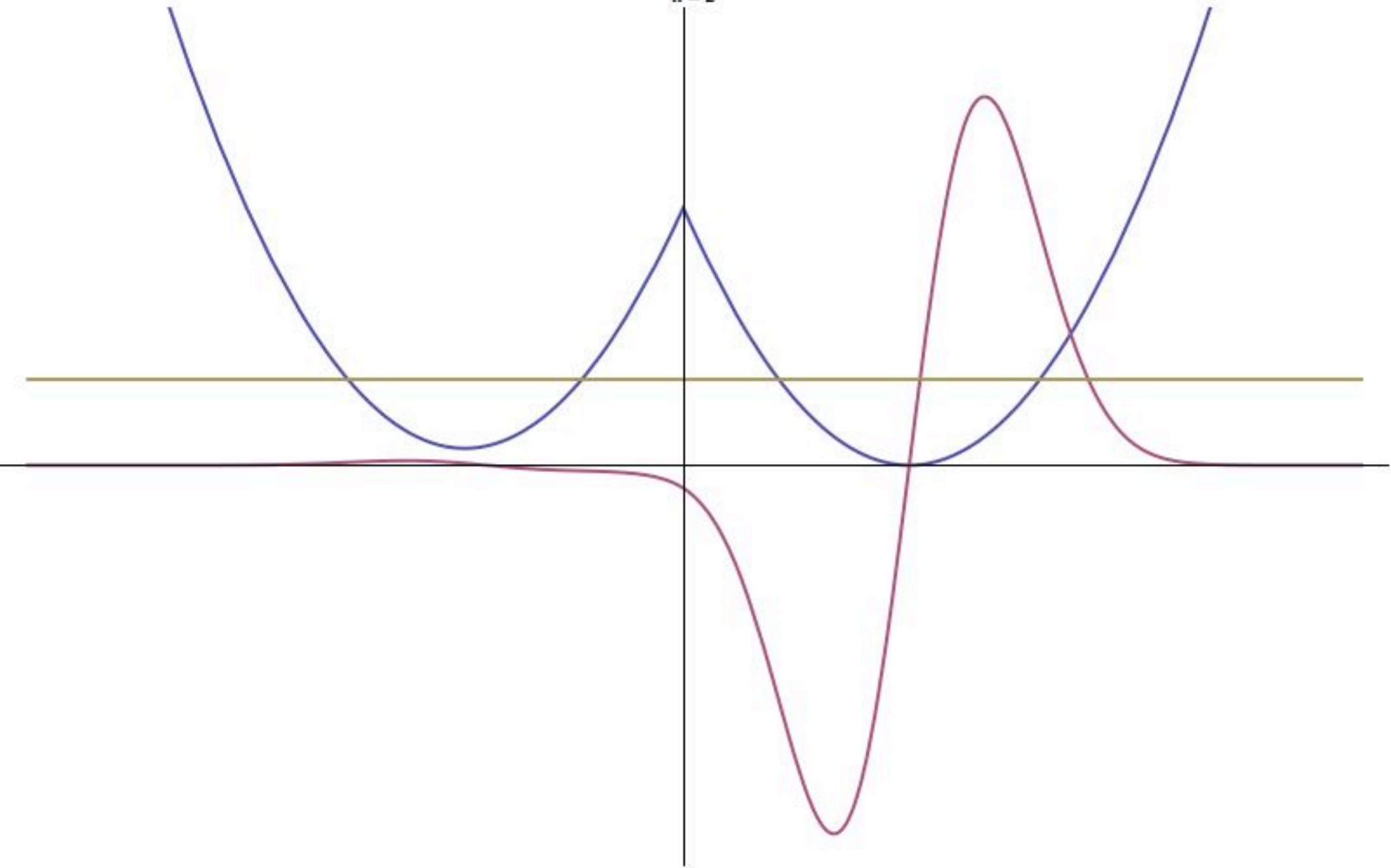


1. Fix 2 boundary conditions here

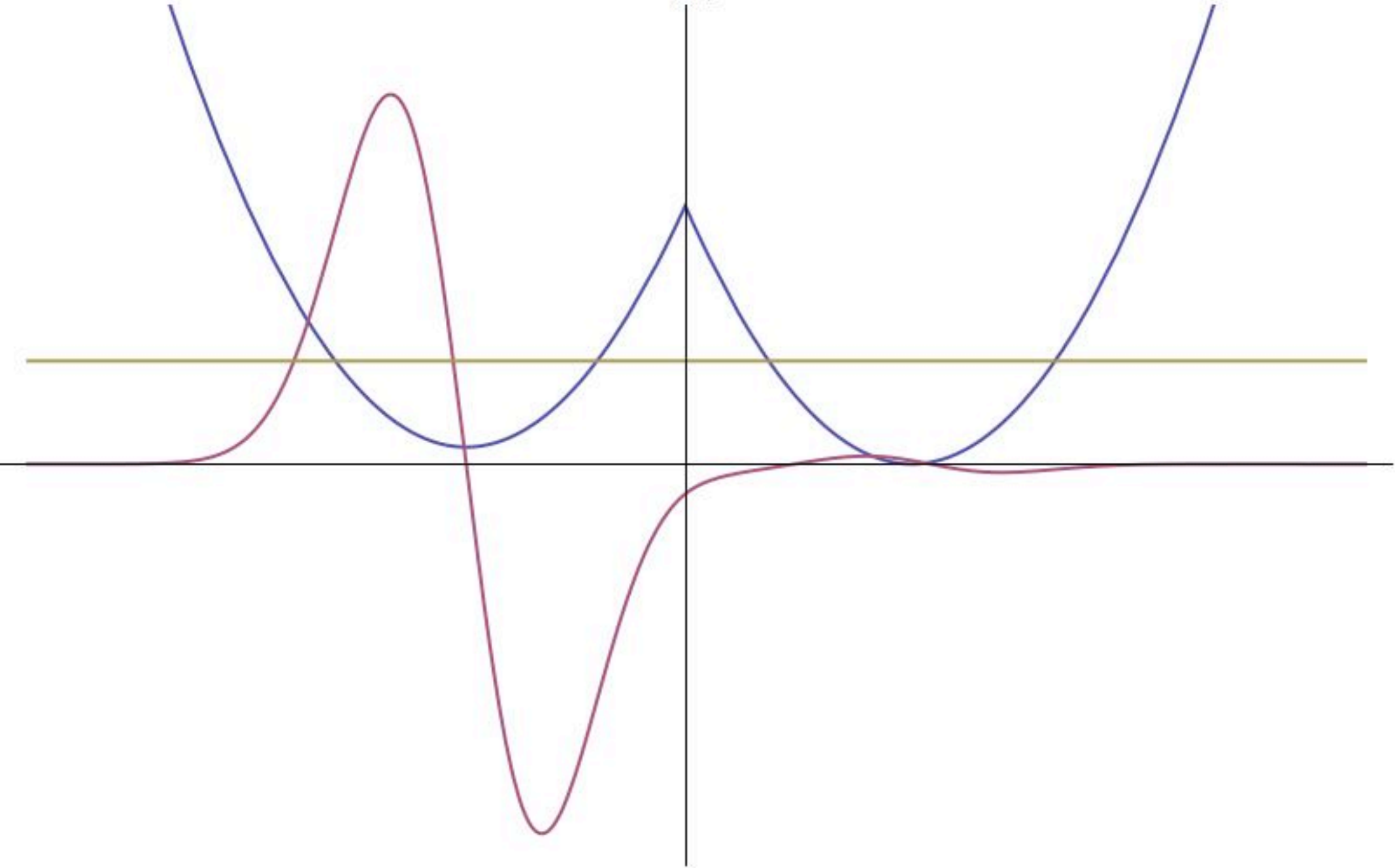
2. Choose energy for convergence here

3. We are left with no free parameters. All wave functions will have a growing and a decaying mode with $\mathcal{O}(1)$ coefficients. One of them will thus always dominate, and all solutions are primarily supported on the left, or the right.

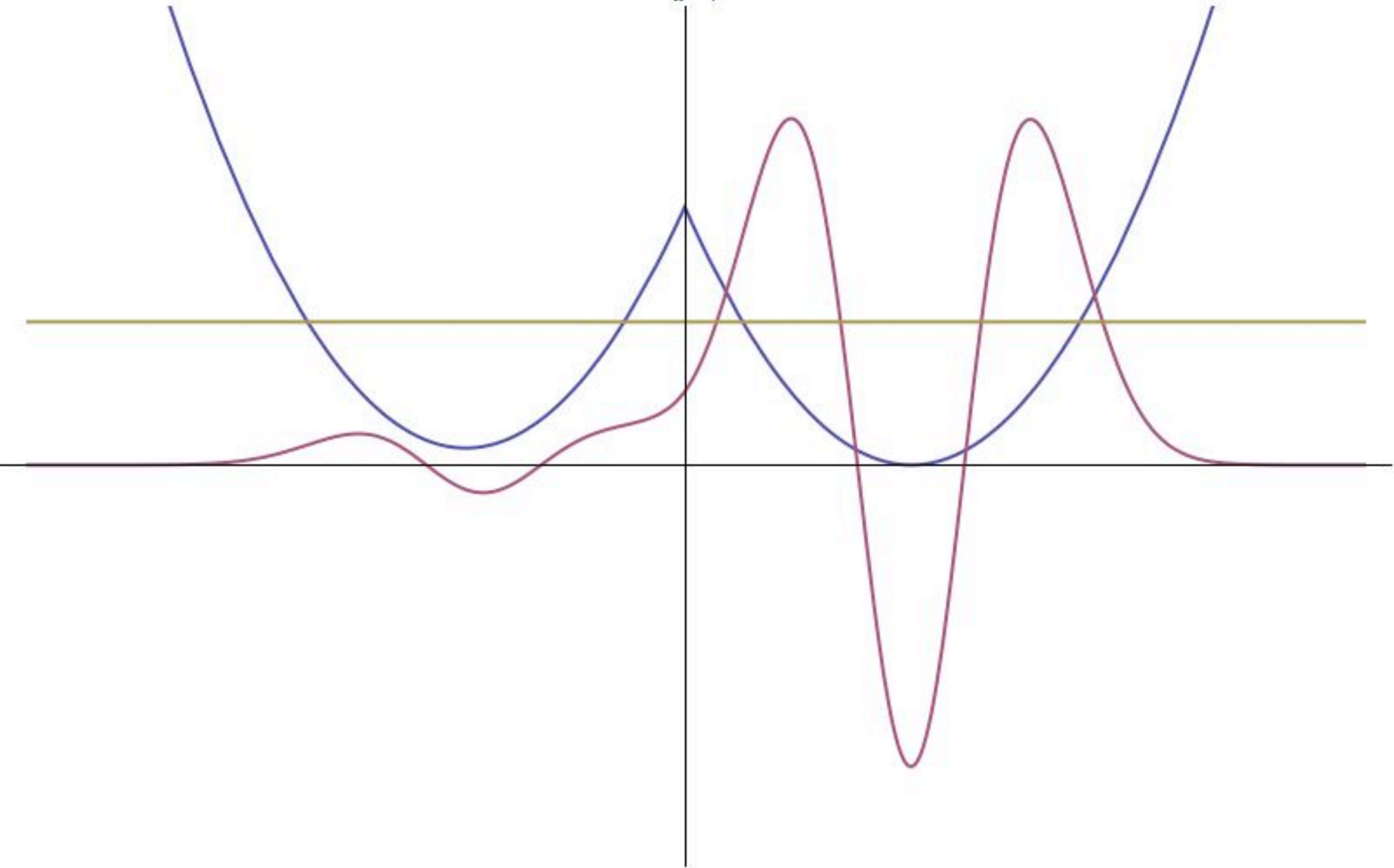
$n = 2$



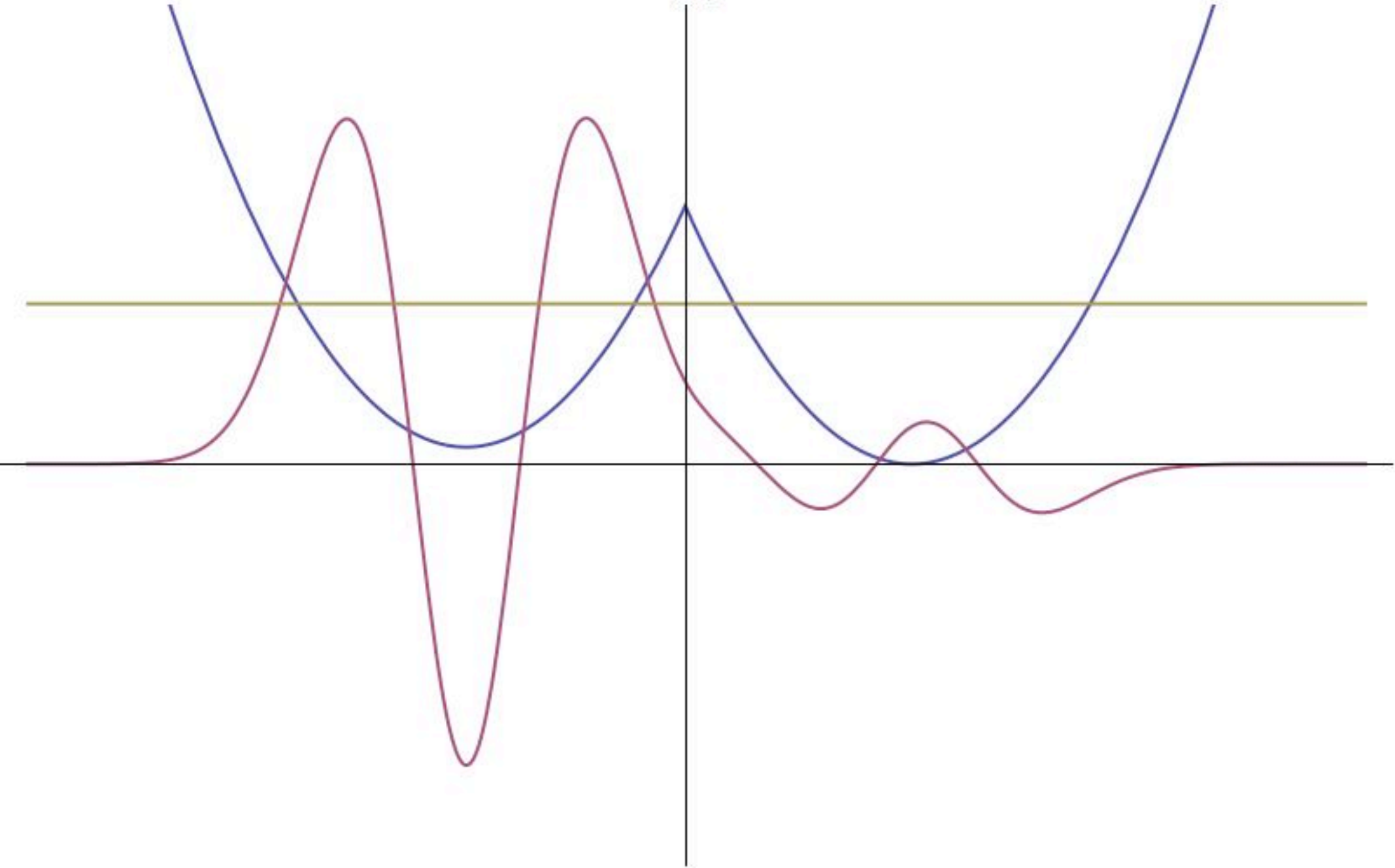
$n = 3$



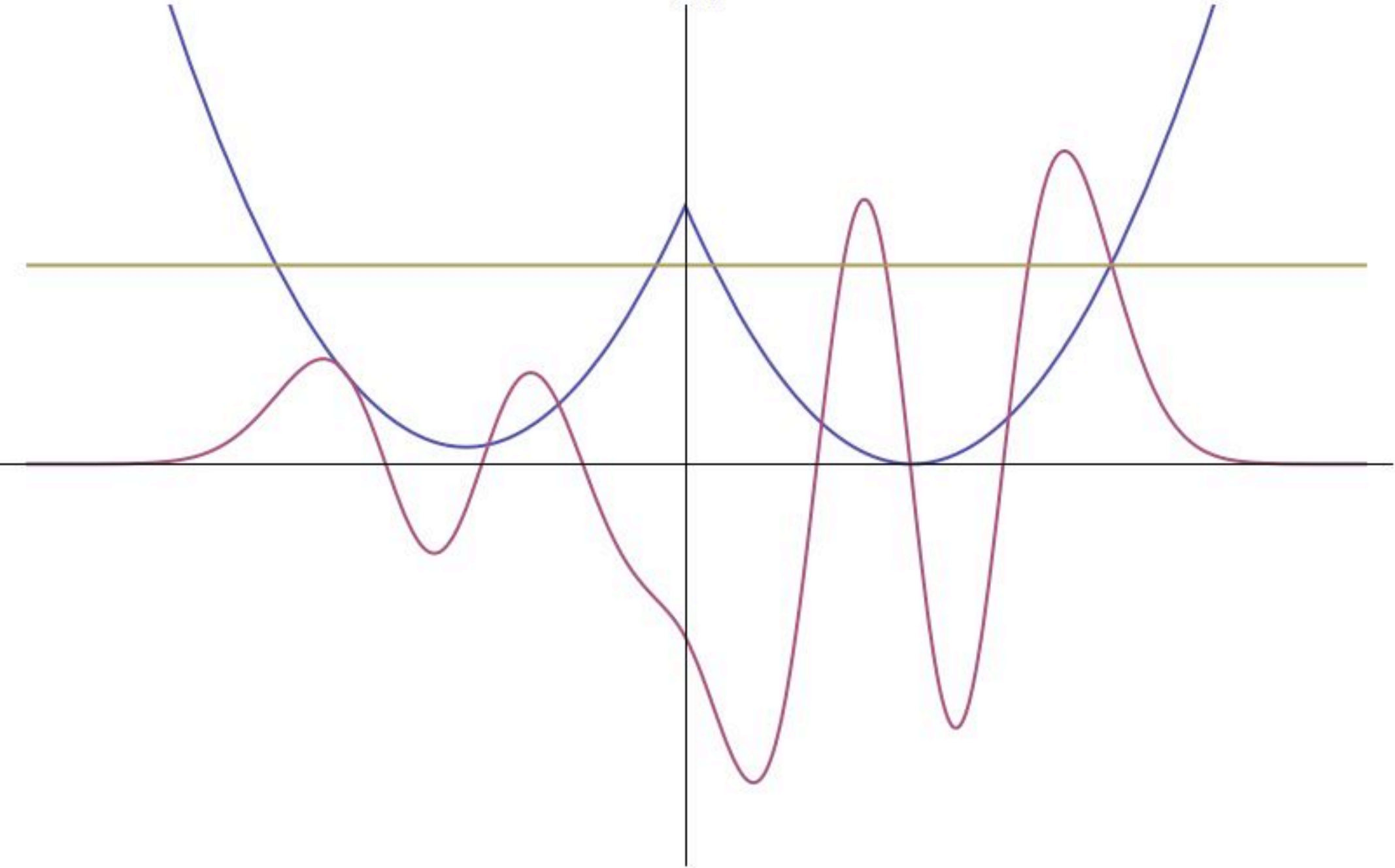
$n = 4$



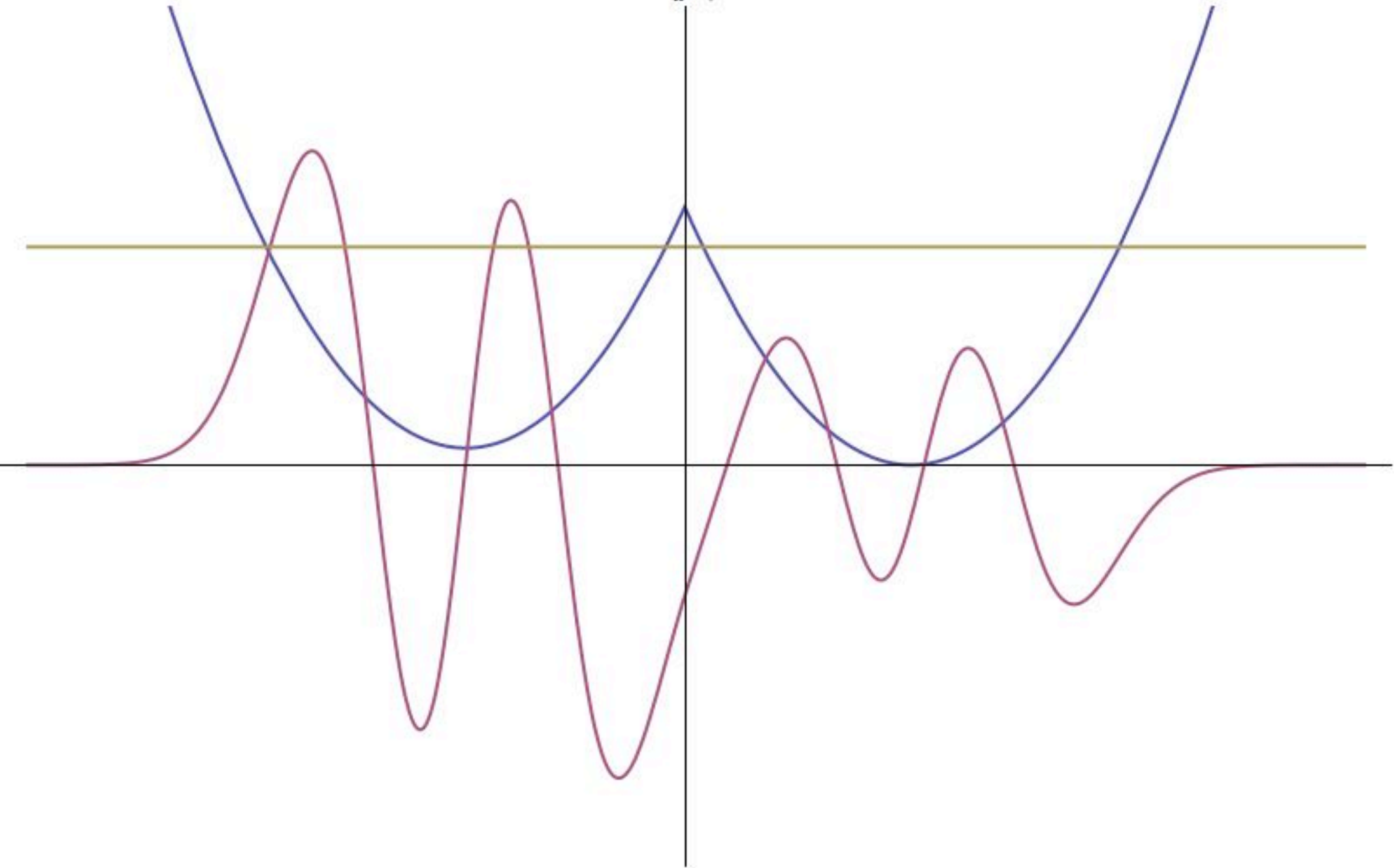
$n = 5$



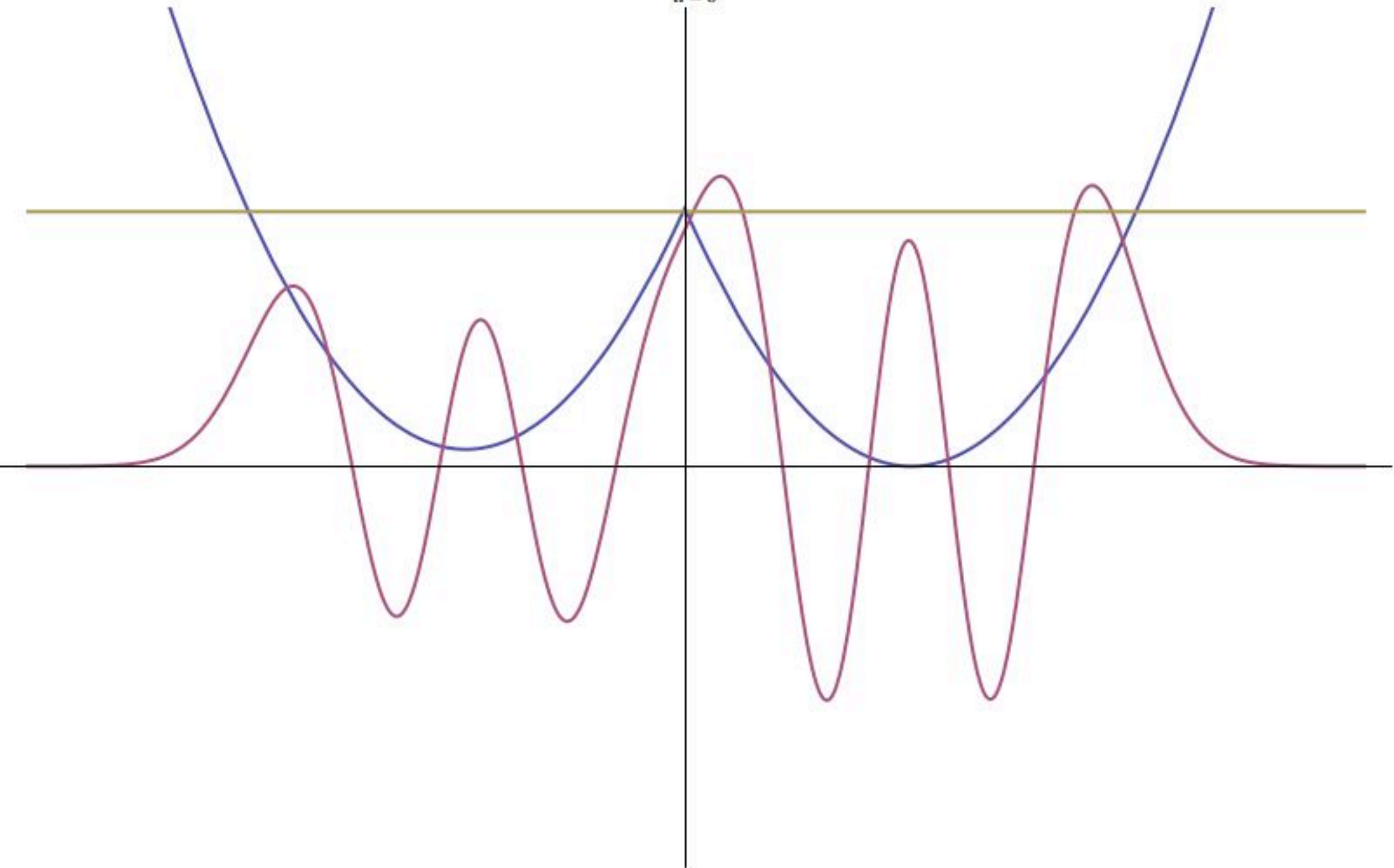
$n = 6$



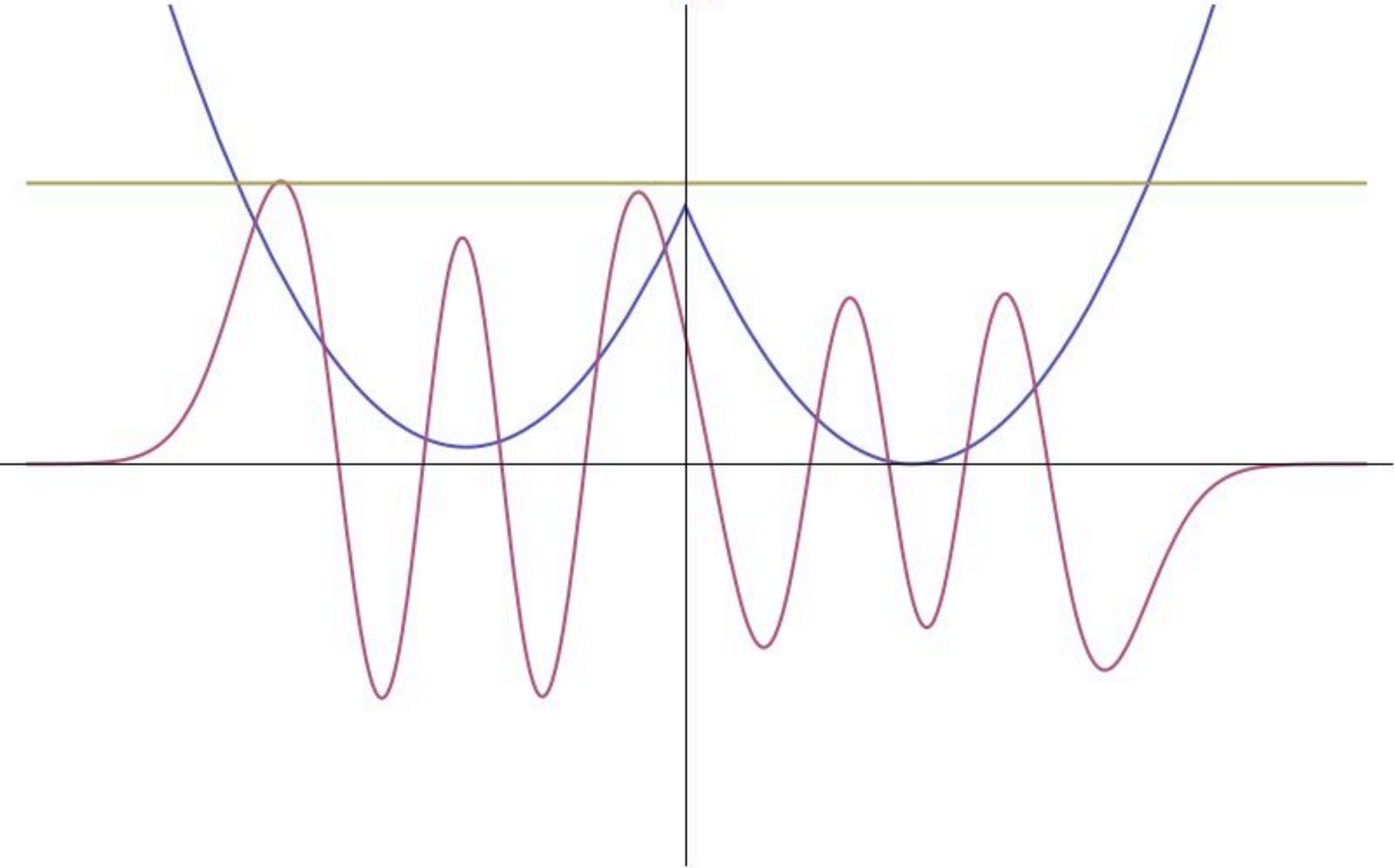
$n = 7$



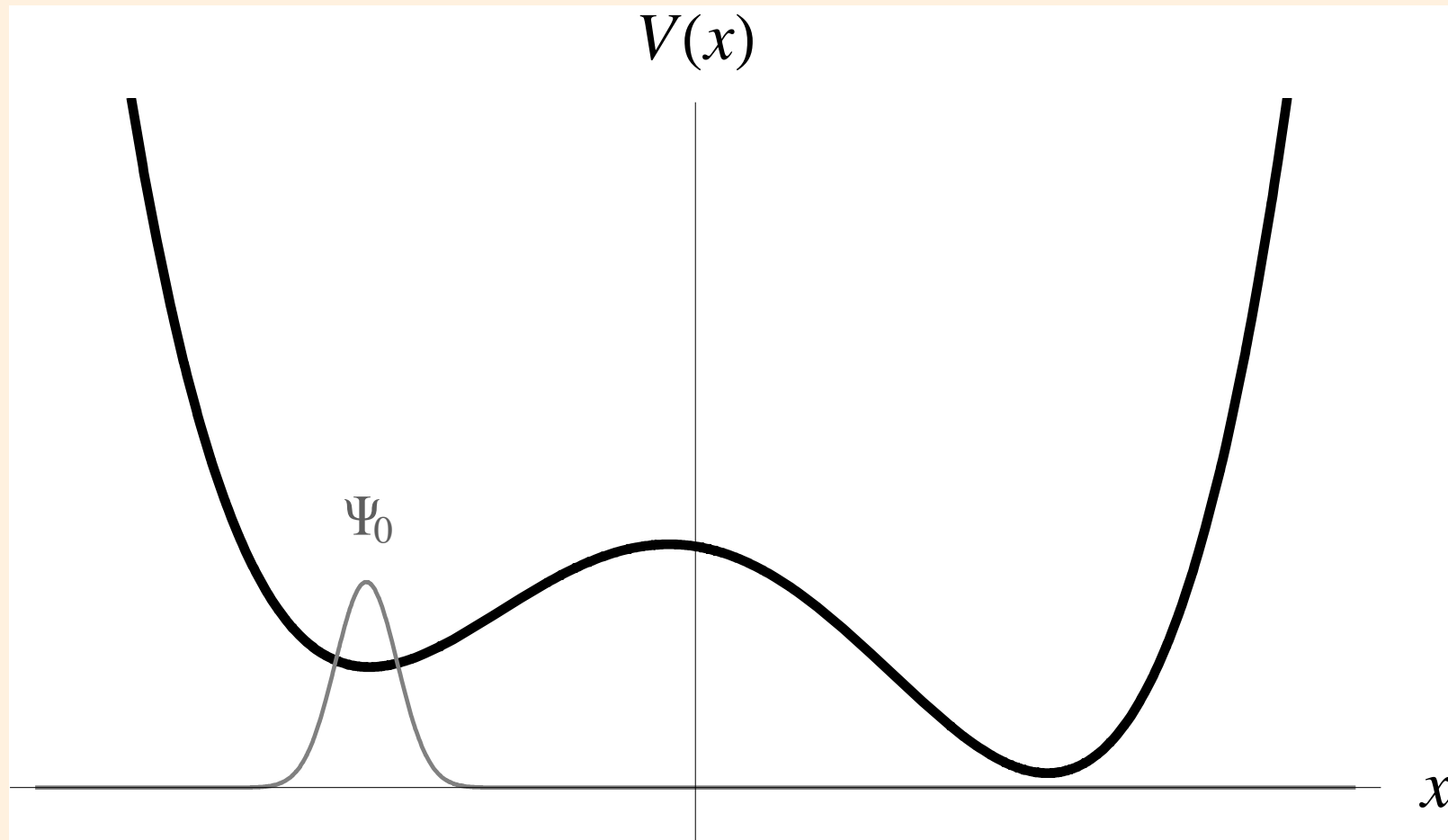
$n = 8$



$n = 9$



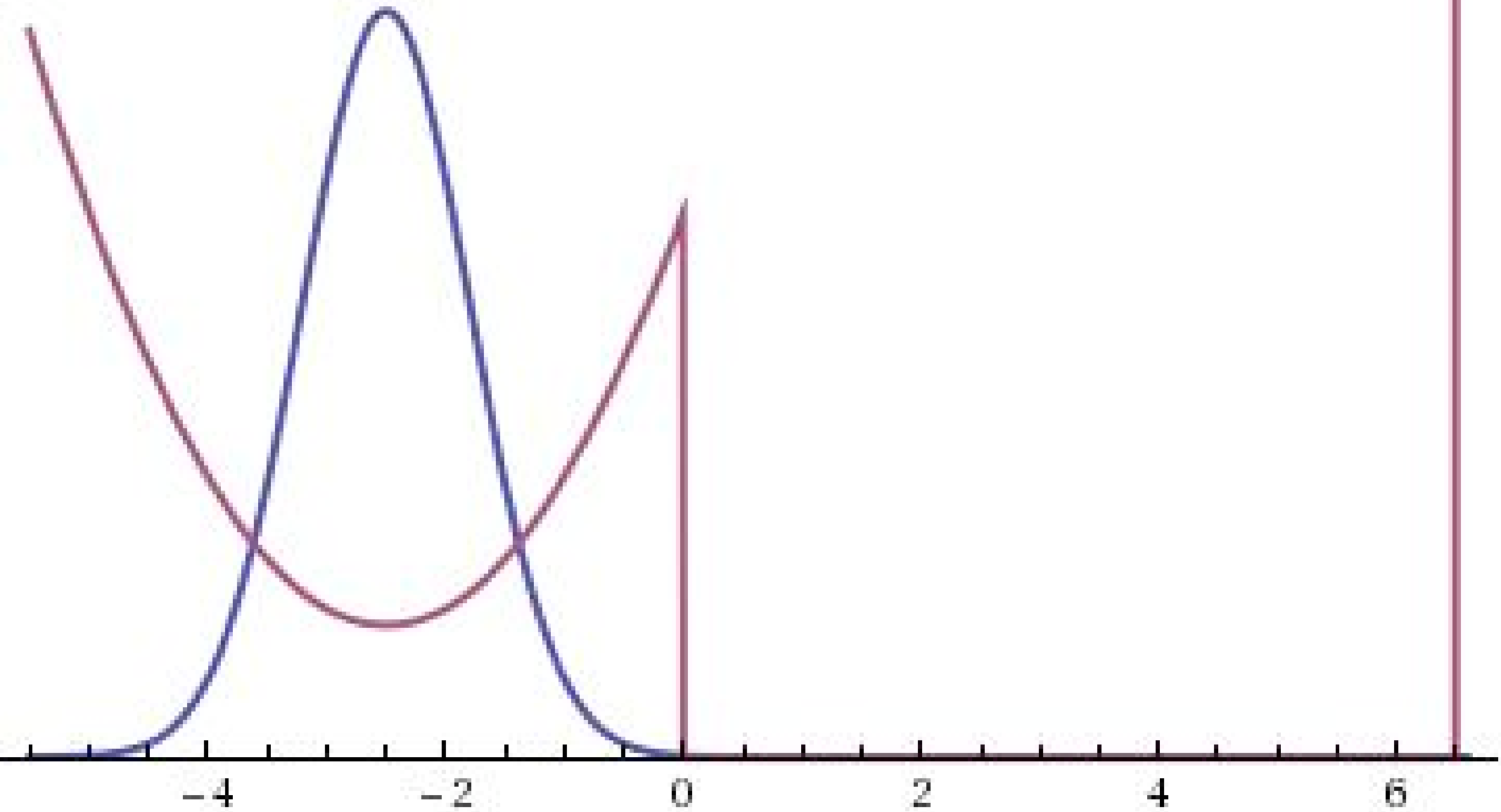
The asymmetric double well has a stable false vacuum



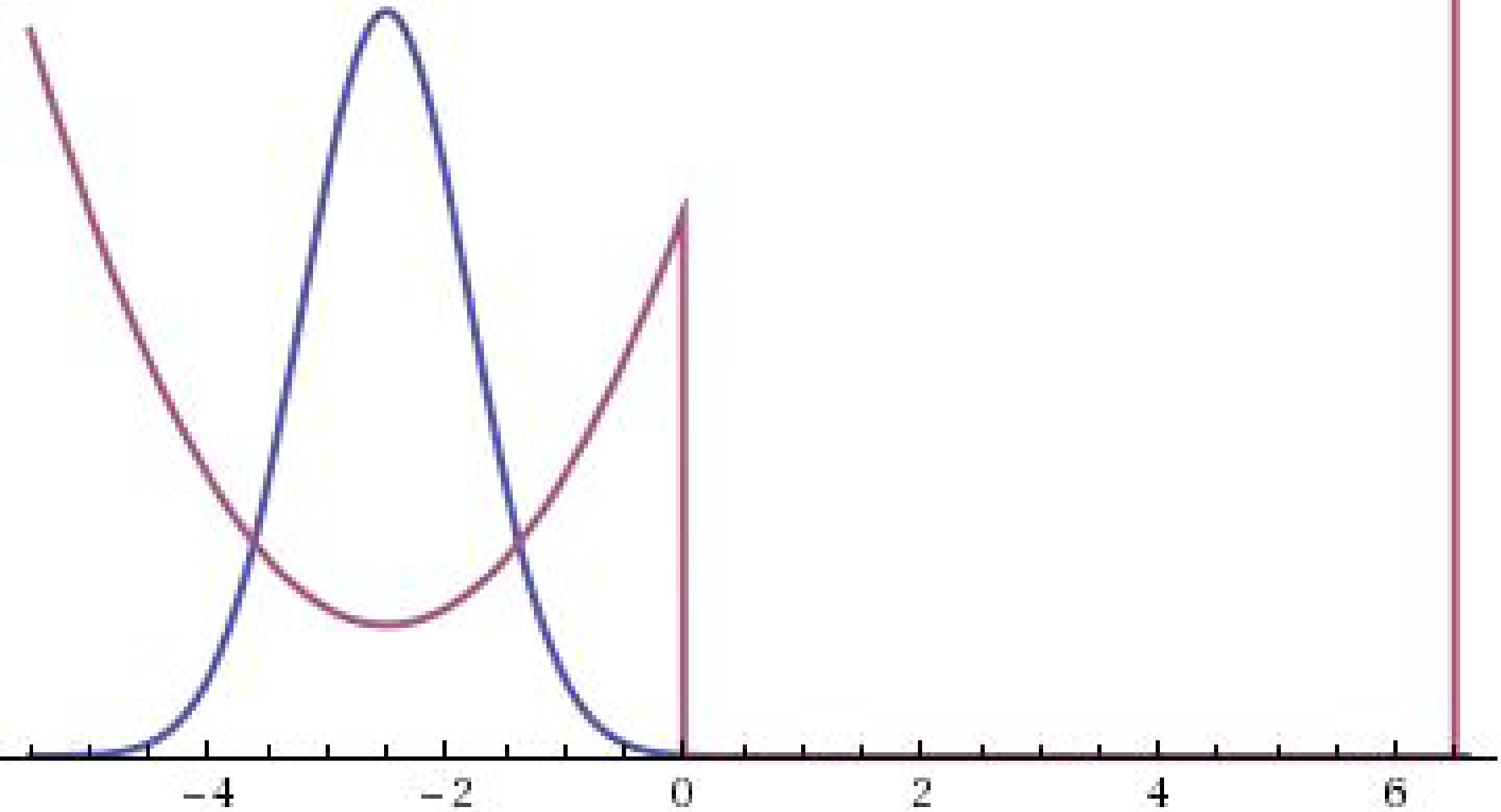
The (LHS) perturbative vacuum **is** an approximate energy eigenstate when $S_B \gg 1$.

$$|\Psi_0\rangle \approx |E_n\rangle$$

$$|\langle x | e^{-iHt} | \Psi_0 \rangle|^2$$

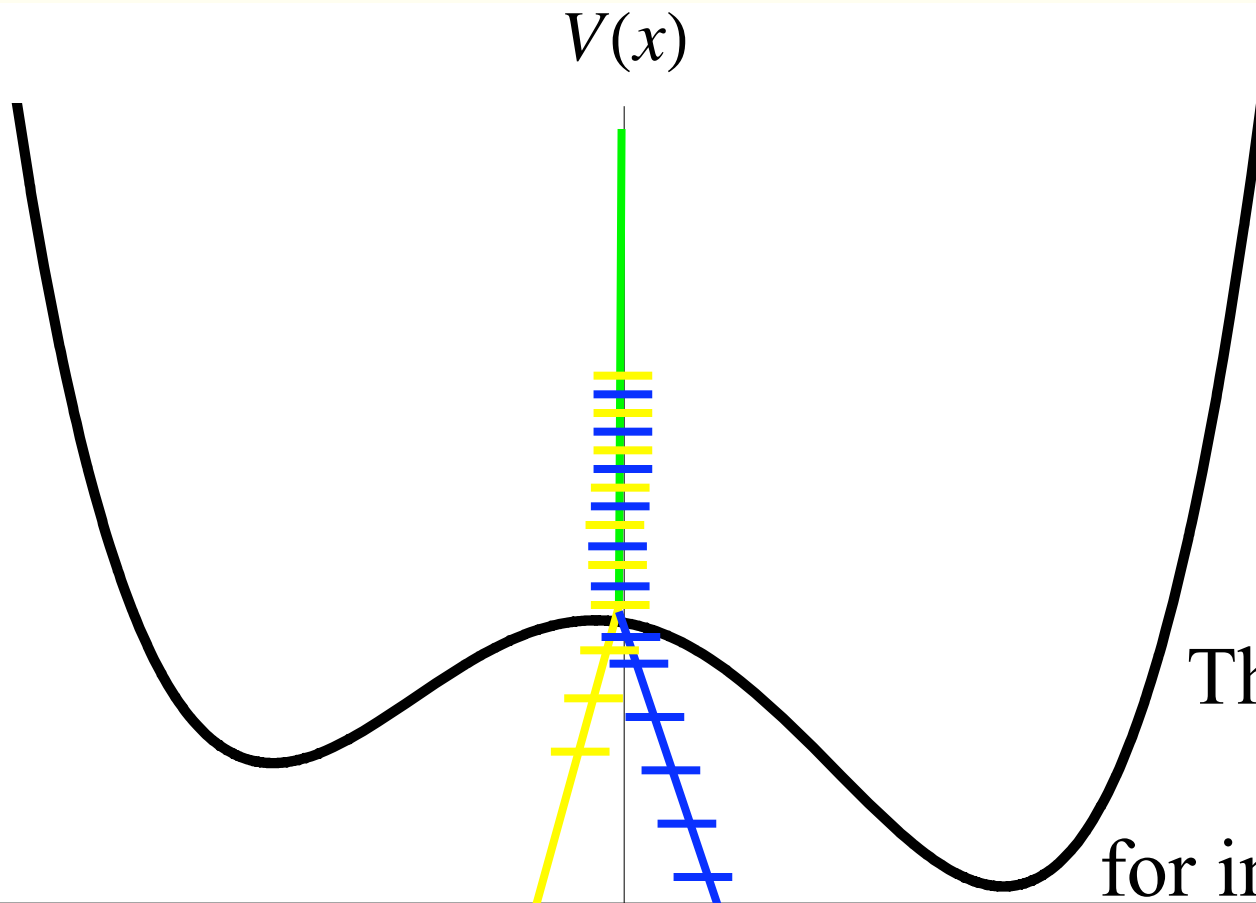


$$|\langle x | e^{-iHt} | \Psi_0 \rangle|^2$$



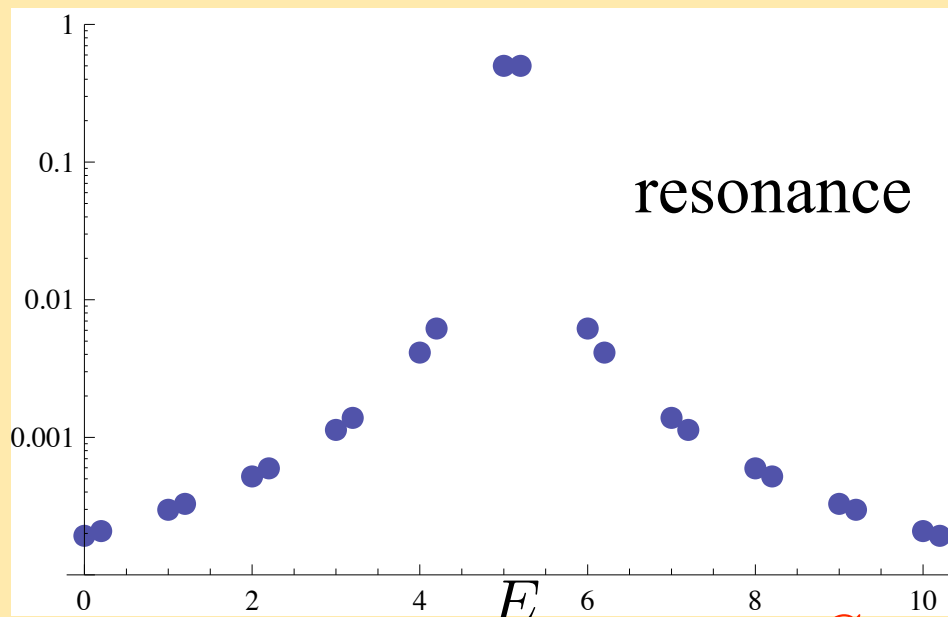
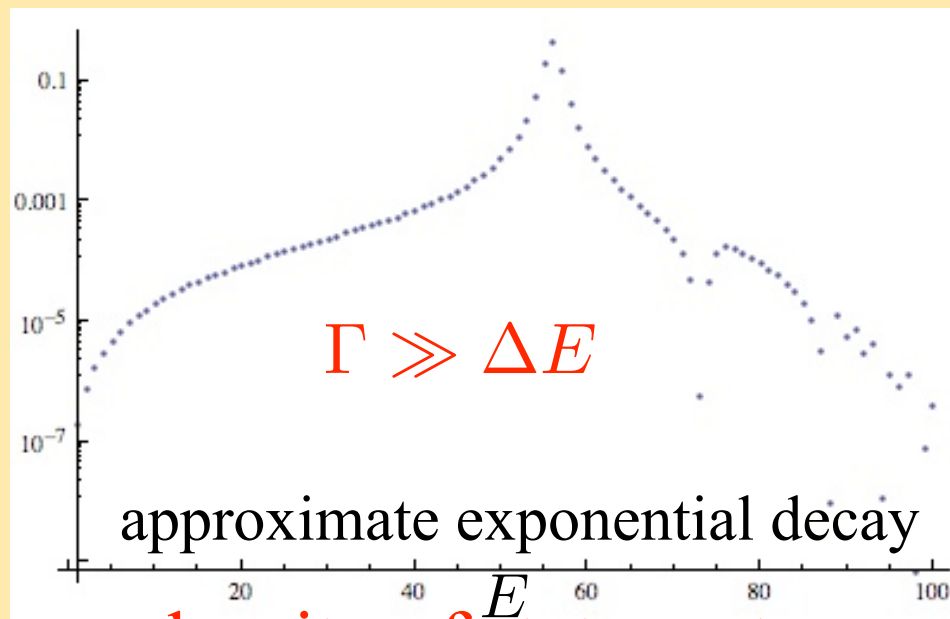
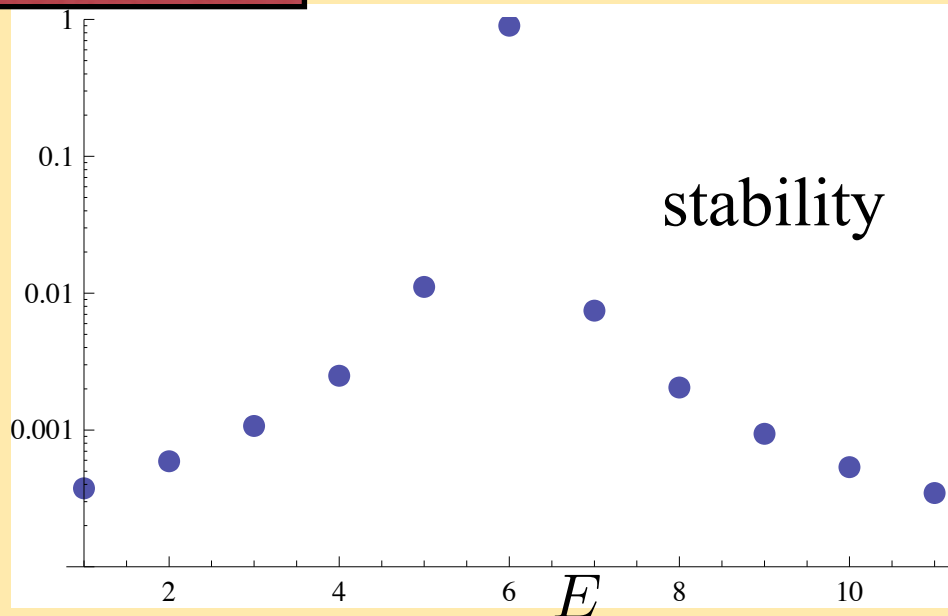
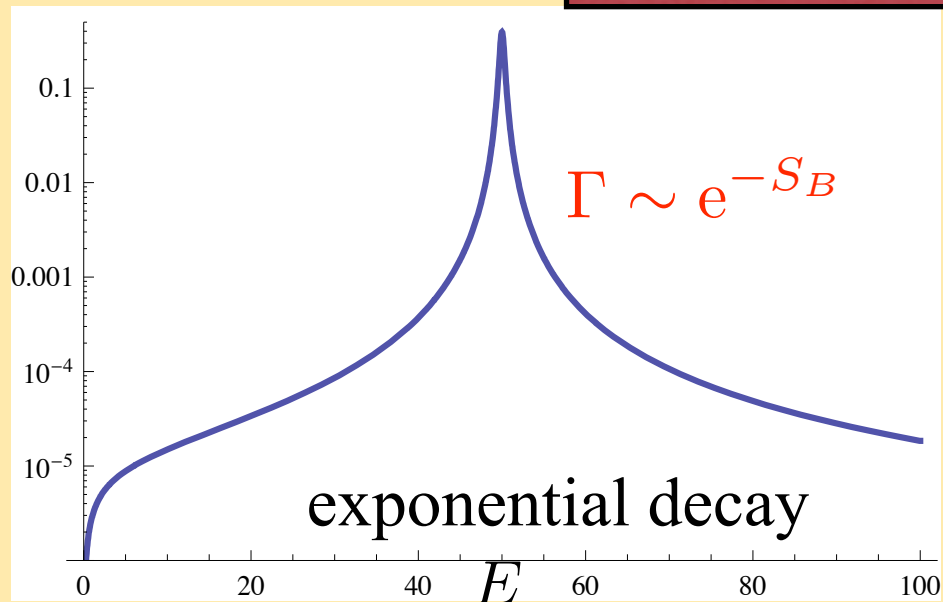
The false vacuum cannot decay because there are no excited true vacuum states with overlapping energy. This is found experimentally in cavity QED.

$$\Delta E_{\Psi_0} \approx \Gamma \sim e^{-S_B}$$
$$\Delta E_{tv} \sim 1/\ell$$



The true vacuum must
be *very* large
for instability to be generic.

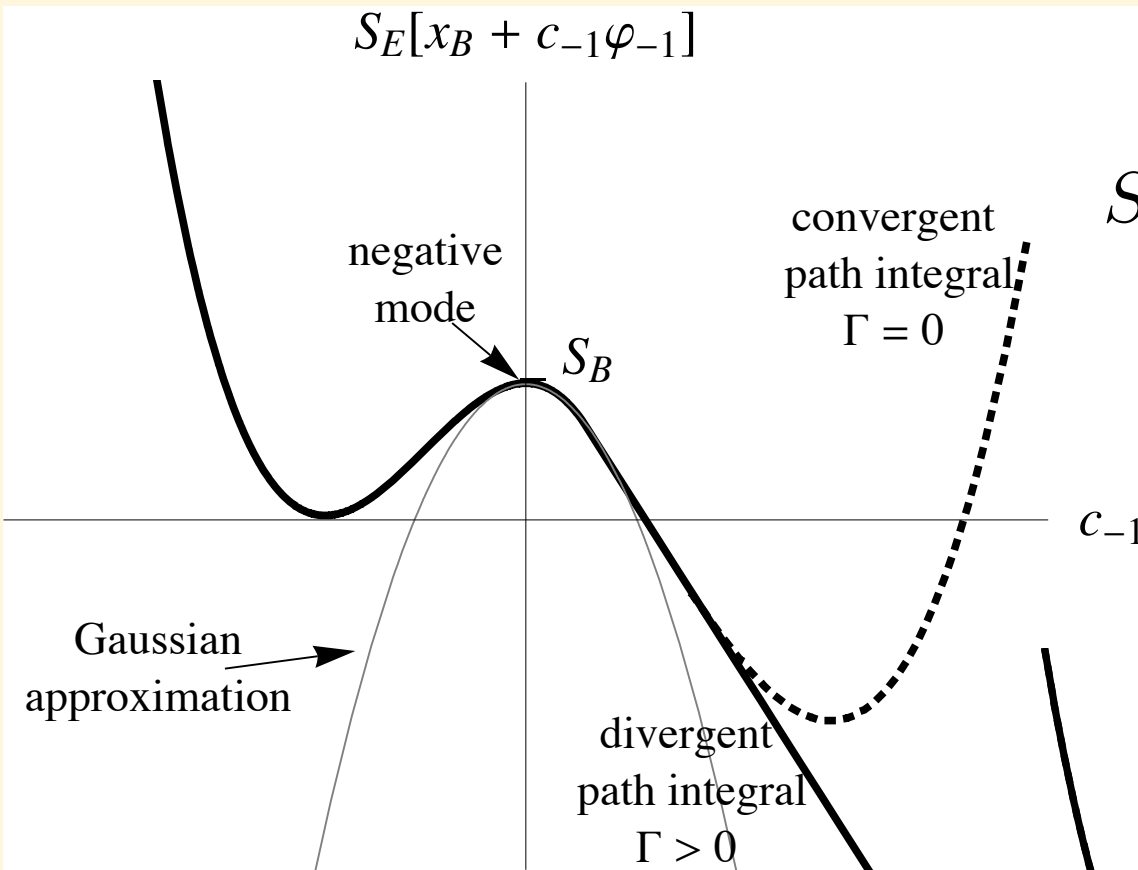
$$|\langle E | \Psi_0(t) \rangle|^2 \leftarrow \text{time independent}$$



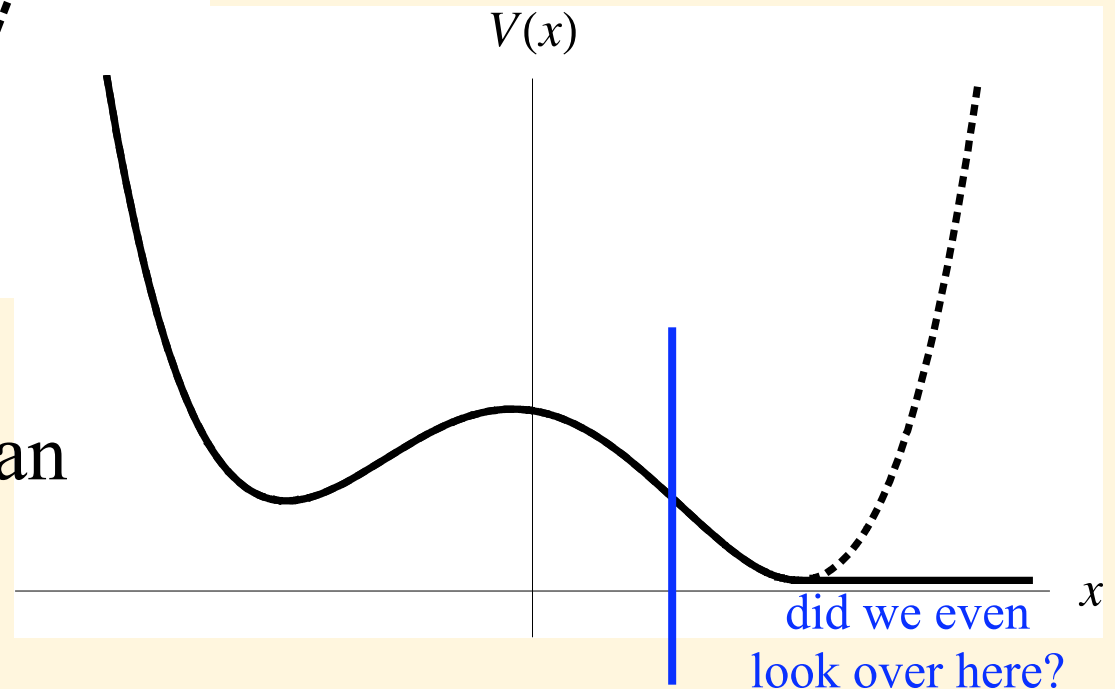
density of states set
by volume

peak width Γ set by e^{-S_B}

The instanton formalism appears to predict exponential decay. The resolution is that the single negative mode is rendered benign by “compactifying” the true vacuum.

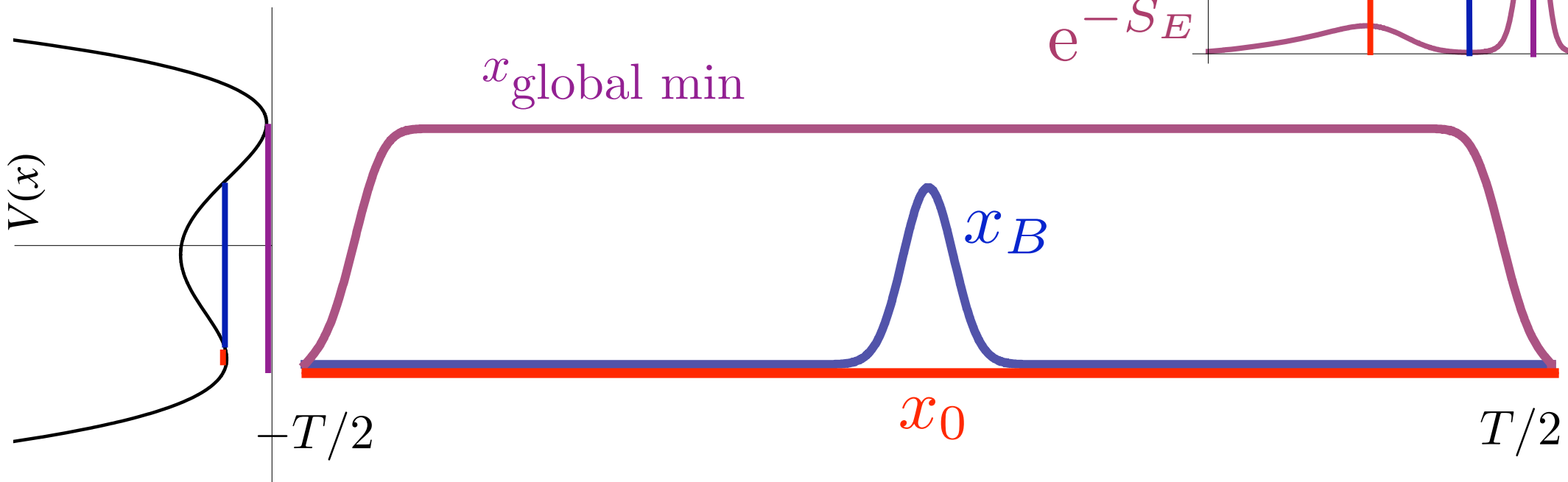


$$S_E = \int_{-T/2}^{T/2} \left(\frac{1}{2} \dot{x}^2 + V(x) \right) d\tau$$



The double well has Euclidean action bounded below!

For the double well, “the bounce” has a *benign* negative mode. The lower action solutions to either side of “the bounce” are both finite. The contribution to the partition function from “the bounce” is negligible.

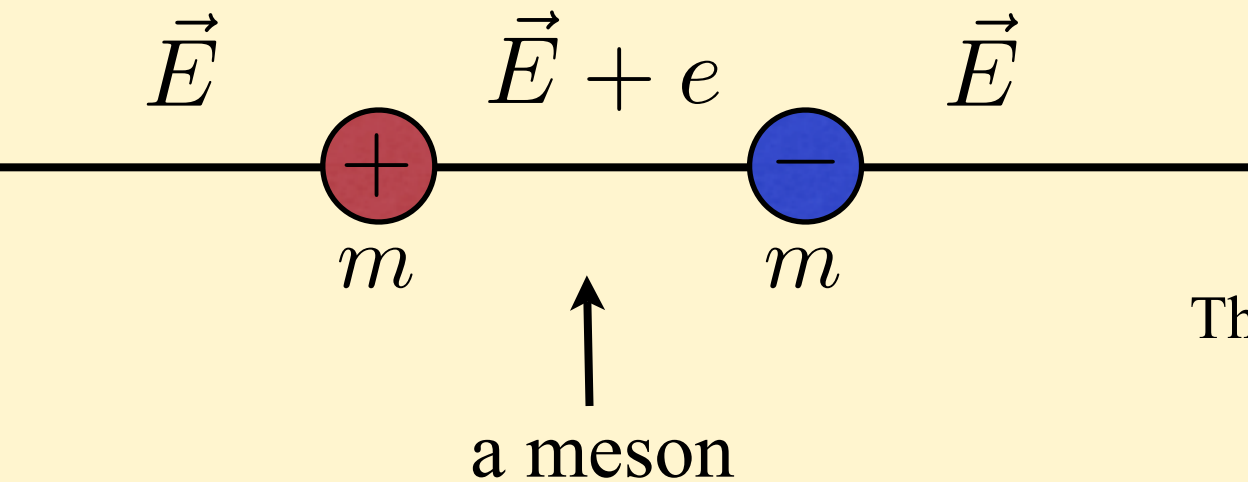


The Schwinger model

vacuum electric
field = $\frac{\theta e}{2\pi}$.

$$S = - \int \left(\frac{1}{4} F^2 + \frac{1}{2} \overline{D_\mu \Phi} D^\mu \Phi + \frac{m^2}{2} \overline{\Phi} \Phi \right) d^2 x + \frac{e}{2\pi} \theta \int F$$

$$\vec{E} = *F \quad (\text{a scalar})$$



There is no photon: low energy mesons are stable

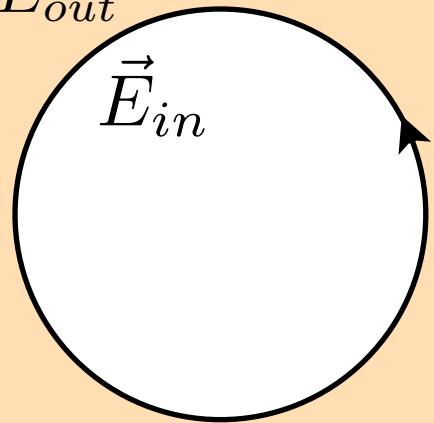
Confinement: All quarks undergo piecewise uniform acceleration

Pair production (vacuum instability)

The Schwinger model $S_E = \oint_{\partial\Sigma} m - \int_{\Sigma} \epsilon$

Exponential decay of the false vacuum is calculated precisely as it was in quantum mechanics:

$\vec{E}_{out} - \vec{E}_{in} = \pm e$ $\mathbb{R}^{1,1} \rightarrow \mathbb{R}^2$ $\epsilon = \frac{1}{2} (\vec{E}_{out}^2 - \vec{E}_{in}^2)$

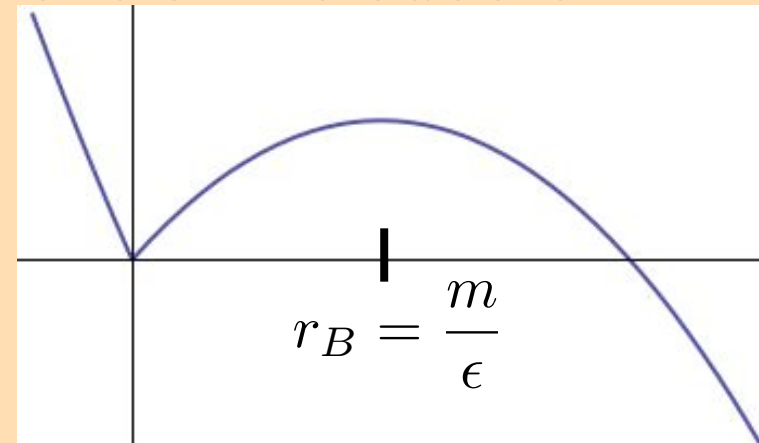


“The Bounce” = S^1 domain wall in \mathbb{R}^2

2 zero modes (translations in \mathbb{R}^2)

Negative mode corresponds to dilatations of the bubble

$$S_E = 2\pi r m - \pi r^2 \epsilon$$



Nucleation rate per unit length

$$\langle \vec{E} | e^{-\mathcal{H}XT} | \vec{E} \rangle = \int [d\Sigma] e^{-S_E[\Sigma]}$$

$$S_E = m \oint_{\partial\Sigma} ds + \frac{\vec{E}^2}{2} \int_{\Sigma} d^2x$$

$$\vec{E}_{in} - \vec{E}_{out} = \pm e$$

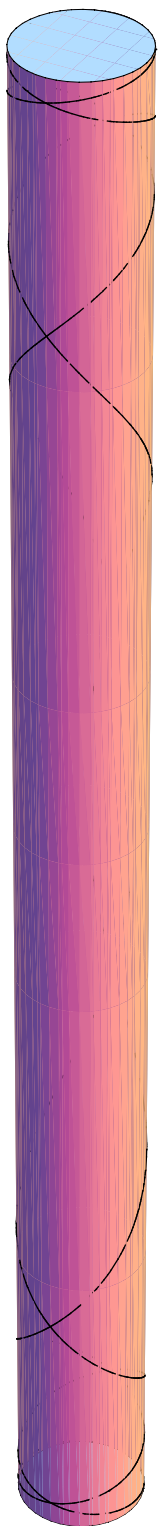
$$\langle \mathcal{H} \rangle = \frac{1}{2} \vec{E}^2 - i \frac{e\vec{E}}{4\pi} \exp\left(-\pi m^2 / e\vec{E}\right)$$

decay rate per unit volume

$$\Gamma = \frac{1}{2} \det' \left[\left| \frac{\delta^2 S_E}{\delta \partial \Sigma^2} \right| \right]^{-\frac{1}{2}} e^{-S_B} = \frac{e\vec{E}}{2\pi} \exp\left(-\pi m^2 / e\vec{E}\right)$$

↑ Sagredo's missing factor

What happens if we compactify the Schwinger model on a circle?



As we dilate the instanton on the cylinder, it will overlap itself.

In the probe approximation, this is irrelevant:

$$S_E = 2\pi r m - \pi r^2 \epsilon \\ \Rightarrow \Gamma > 0$$

The negative mode is still “dangerous”

This can be confirmed by the method of Bogolyubov coefficients.

But if we include the back-reaction on the electric field, the action is bounded below.

The negative mode is actually “benign”

$$\Rightarrow \Gamma = 0$$

This can be confirmed by explicit computation of the discrete spectrum when $\Gamma l \ll 1/l$

3+1 d QFT in flat space

$$S_E = \int_{\mathbb{R}^4} \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + V(\phi) d^4x$$

Integrate out the UV:
only domain walls remain

$$S_E = \mu \int_{S^3} d^3x - \epsilon \int_{B^4} d^4x$$

a.k.a. Brown - Teitelboim

Exponential decay of the false vacuum is calculated precisely as it was in the Schwinger model:

$$\mathbb{R}^{1,3} \rightarrow \mathbb{R}^4 \quad ds^2 \rightarrow d\tau^2 + dx^2 + dy^2 + dz^2$$

“The Bounce” = S^3 domain wall in \mathbb{R}^4

4 zero modes (translations in \mathbb{R}^4)

Negative mode corresponds to dilatations of the bubble

It is “dangerous.”

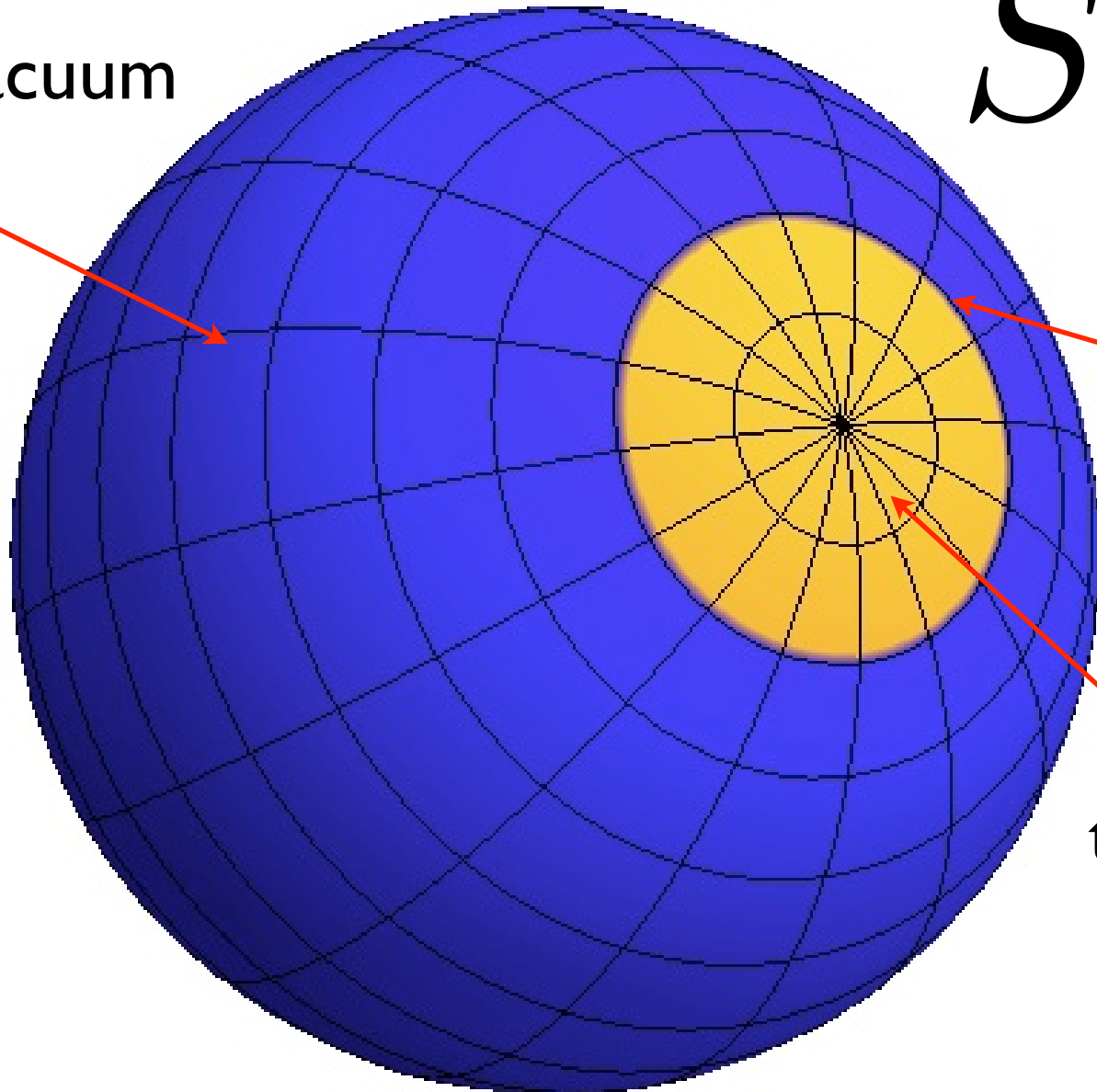
$$\Gamma \sim e^{-S_B} \quad S_B = \frac{27\pi^2 \mu^4}{2\epsilon^3}$$

QFT in de Sitter space

$$ds^2 = \left(1 - \frac{r^2}{l^2}\right) d\tau^2 + \left(1 - \frac{r^2}{l^2}\right)^{-1} (dr^2 + r^2 d\Omega^2)$$

false vacuum

B^4



S^4

see, however, recent work by Polyakov

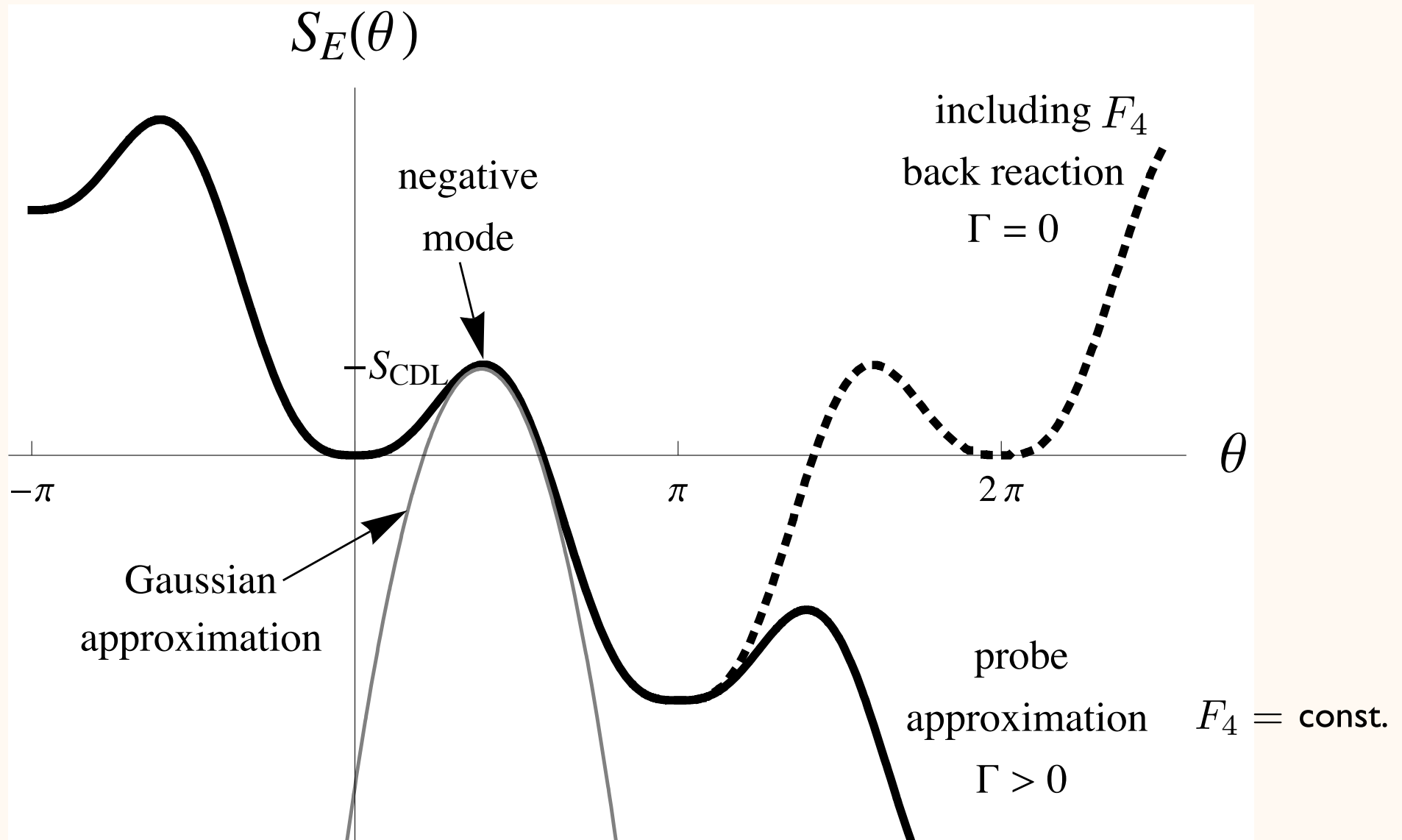
S^3

bubble wall

true vacuum

B^4

Coleman - De Luccia action as a function of bubble radius



The Coleman - De Luccia negative mode is *benign* in de Sitter space.

Caveats

De Sitter space is not a box: there is certainly a continuous spectrum, so our intuition about finite volumes is not useful. $M_{Pl} \rightarrow \infty$

Just because there is *a* stable de Sitter false vacuum doesn't mean it is the *relevant* one (for, say, eternal inflation). I have argued in favor of the existence of a de Sitter invariant false vacuum. (There is no Poincaré invariant false vacuum in flat space.)

Formalism predicts no exponential decay in dS for any parameter range. We can be sure this breaks down at low curvature just like finite volume QFT, but explicit verification is difficult.

What about the Gibbons Hawking temperature? Is that an *external* heat bath, making QFT in de Sitter space (static patch) non-unitary?

Throughout my analysis I assume GH entropy is entanglement entropy, not an external heat bath.

A toy model...

An example where back-reaction can be ignored: semi-bounded $1+1d$ Rindler space with a uniform electric field.

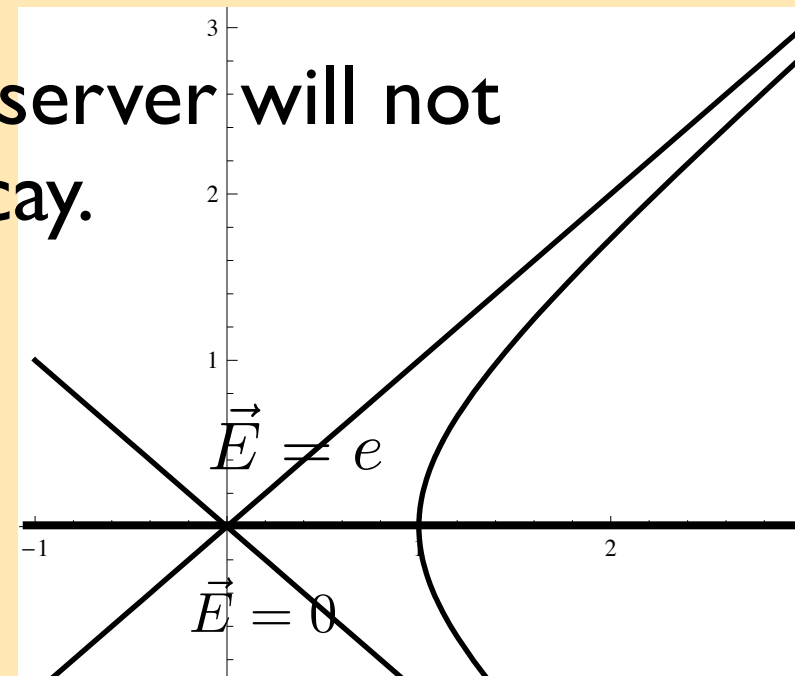
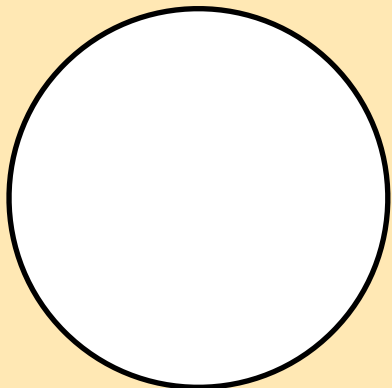
no firm boundary between:

Virtual pair production \longleftrightarrow thermal activation \longleftrightarrow quantum tunneling

by an argument similar to Brown & Weinberg

Schwinger pairs are produced *at rest* w.r.t. the initial conditions (instantaneously generated electric field).

Even a low acceleration Rindler observer will not experience vacuum decay.



Vacuum instability

vs.

Black hole instability

an analogy

Flat

Anti- de Sitter

De Sitter

<p>False vacuum is always unstable (neglecting gravitational back-reaction) Coleman</p>	<p>False vacuum unstable if energy difference is large enough Coleman - De Luccia</p>	<p>False vacuum is stable.</p>
<p>Black holes eat up space at any nonzero temperature. Gross - Perry - Yaffe</p>	<p>Black hole instability only if temperature is large enough Hawking - Page</p>	<p>Black holes are insignificant fluctuations Ginsparg - Perry</p>
<p>unstable</p>	<p>depends</p>	<p>stable</p>

Conclusion

- At semi-classical level, de Sitter space is stable*
- CDL instanton does not mediate decay in dS
- Method of Bogolyubov coefficients does not predict decay unless backreaction is ignored.
- de Sitter space is known to be stable against black hole nucleation
- This is not a question of quantum gravity, which we ignore
- The global picture of the Landscape is complex, but it is far from clear that the Coleman - De Luccia instanton is applicable.

* I assume QFT in de Sitter space is unitary, i.e., there is no external heat bath.